

“Computational Pattern Analysis in Physics and Cosmology”

Reinaldo R. Rosa¹,

Murilo S. Dantas, Gustavo Zaniboni¹, Cristiane P. Camilo¹,
Ana Paula Andrade², André L. Ribeiro², German Gomero³,
William S. Hipólito-Ricaldi³, Carlos Alexandre Wuensche⁴

Haroldo F. C. Velho, Fernando M. Ramos¹,
Cesar A. Caretta¹, Nei Cardenuto¹

¹Núcleo para Simulação e Análise de Sistemas Complexos (NUSASC)
Laboratório Associado de Computação e Matemática Aplicada (LAC)
Instituto Nacional de Pesquisas Espaciais (INPE)
São José dos Campos – SP-Brasil

²UESC, ³IFT/CBF, ⁴DAS/INPE

Versão atualizada do trabalho apresentado no evento
“Conference on Computational Cosmology”
Abdus-Salam ICTP-Trieste, Maio de 2005

Classification of Physical Patterns (Representation Level):

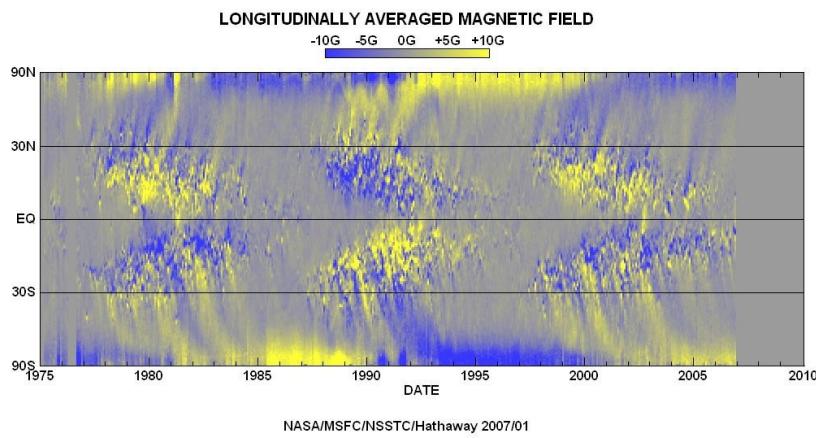
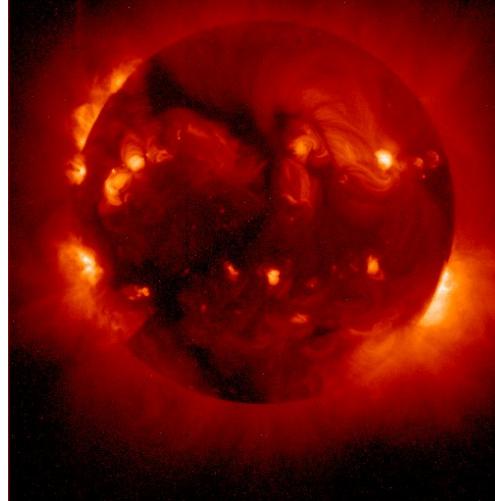
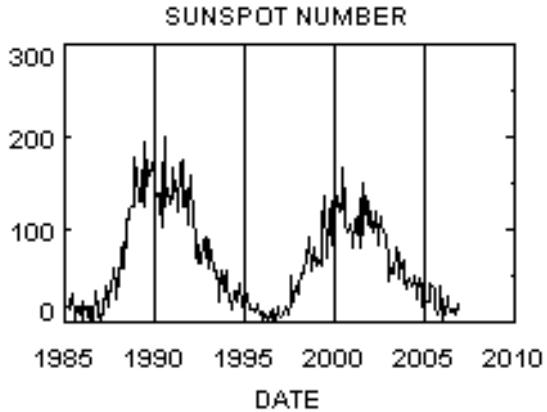
1st Order Patterns (Usual domains): time, space and frequency

2nd Order Patterns: phase space

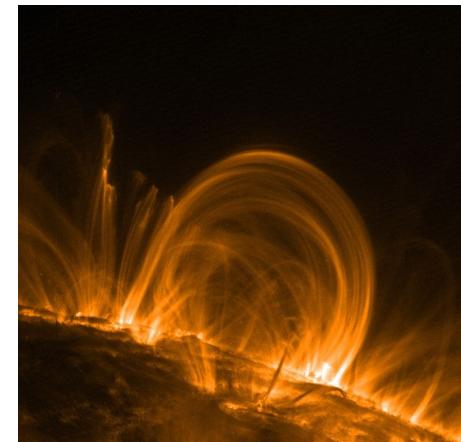
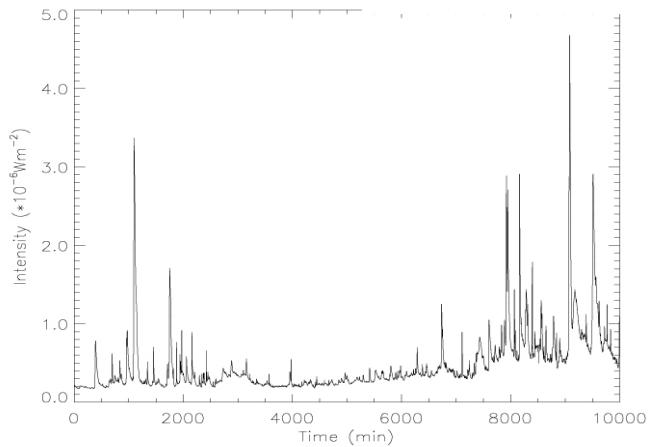
3rd Order (Statistical Parameters): Histogram and Power-law

→ Quantifiers and Metrics → Phenomenological Characterization

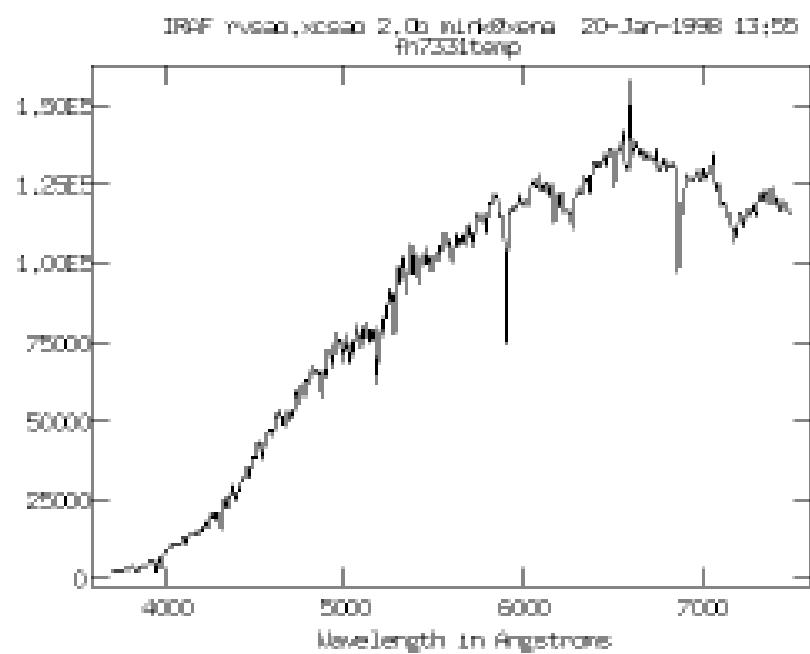
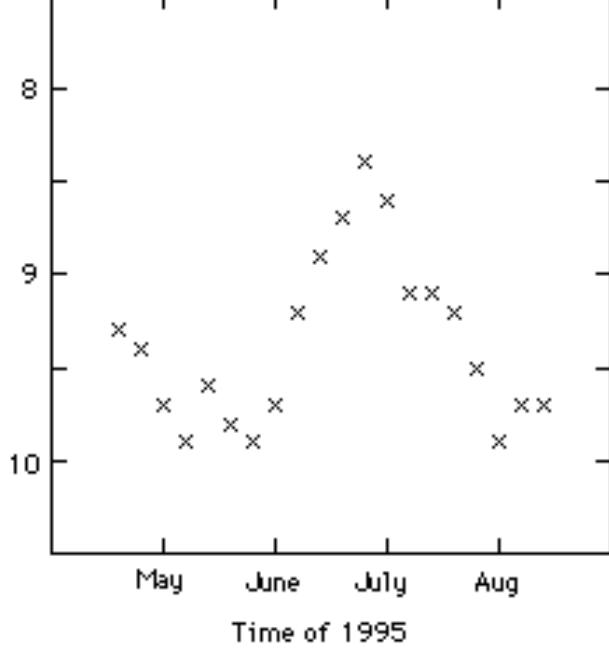
(mencionar noção de vetor de atributos e mineração de dados!)

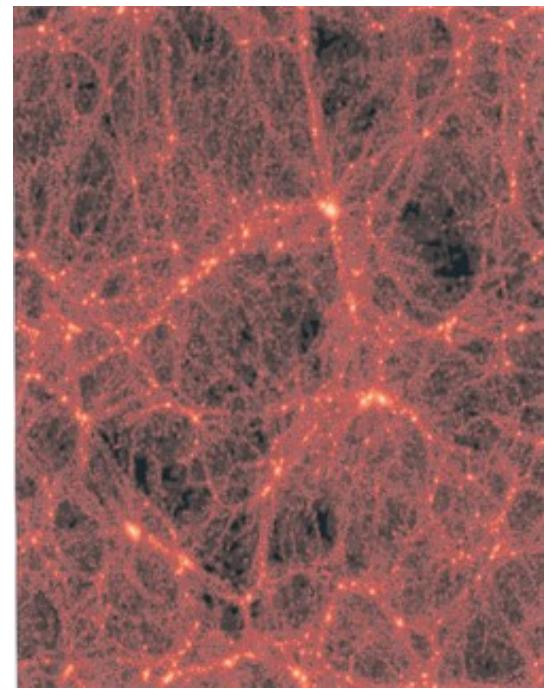
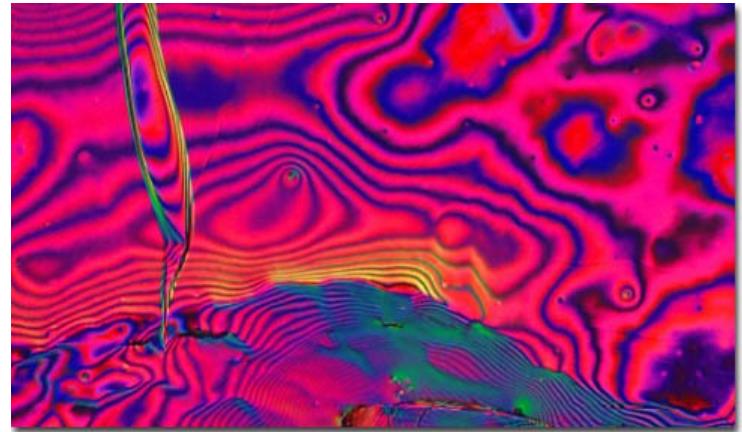


NASA/MSFC/NSSTC/Hathaway 2007/01



Brightness (Magnitude)





At one point, Edward Lorenz was looking for a way to model the action of the chaotic behavior of the gaseous system first mentioned above. Lorenz took a few "Navier-Stokes" equations, from the physics field of fluid dynamics. He simplified them and got as a result the following three-dimensional system:

$$dx/dt = \delta * (y - x)$$

$$dy/dt = r * x - y - x * z$$

$$dz/dt = x * y - b * z$$

Here δ represents the "Prandtl number." This number, which one absolutely does not have to know the meaning of, is

the ratio of the fluid viscosity of a substance to its thermal conductivity (named after Ludwig Prandtl, a German physicist).

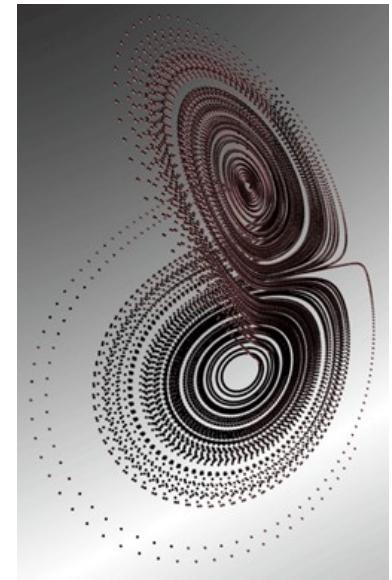
The value Lorenz used was 10.

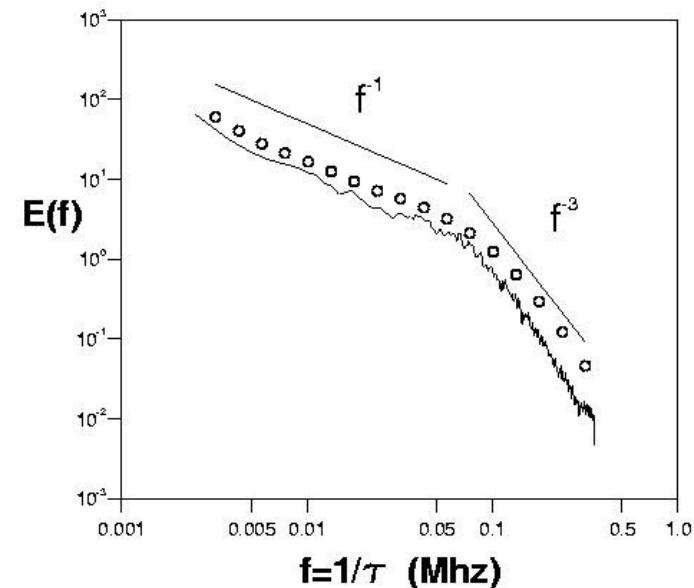
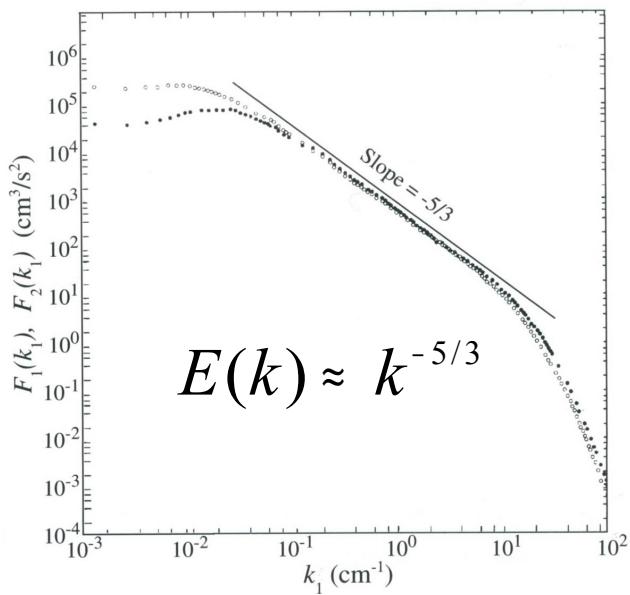
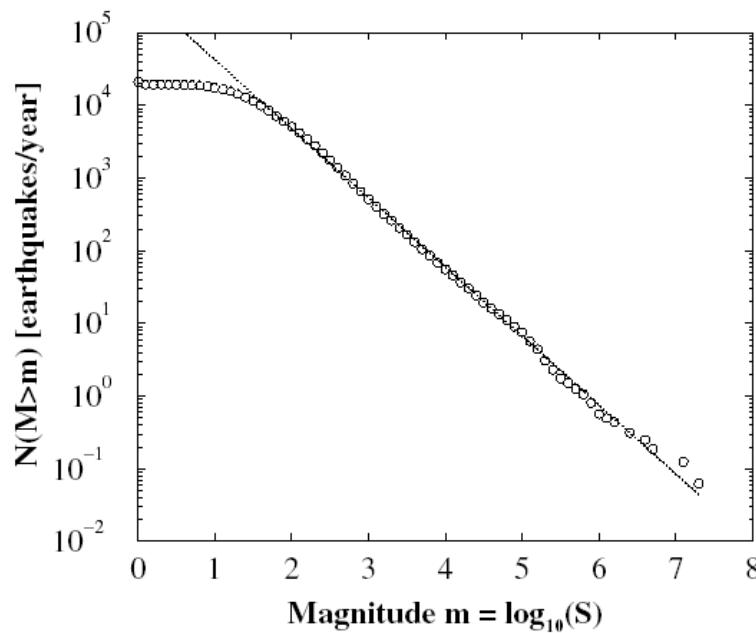
The variable r represents the difference in temperature between the top and bottom of the gaseous system. The value usually used in sample Lorenz attractors such as the one displayed here is 28.

The variable b is the width to height ratio of the box which is being used to hold the gas in the gaseous system.

Lorenz happened to choose $8/3$, which is now the most common number used to draw the attractor.

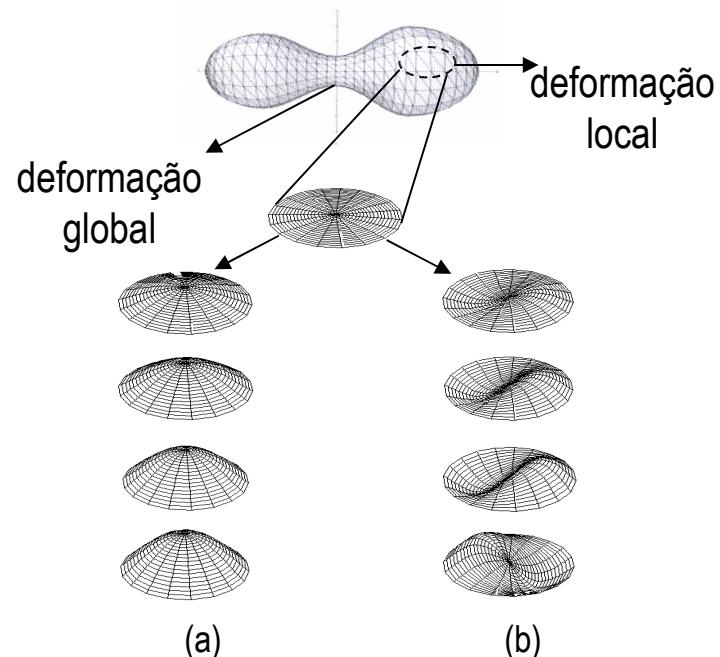
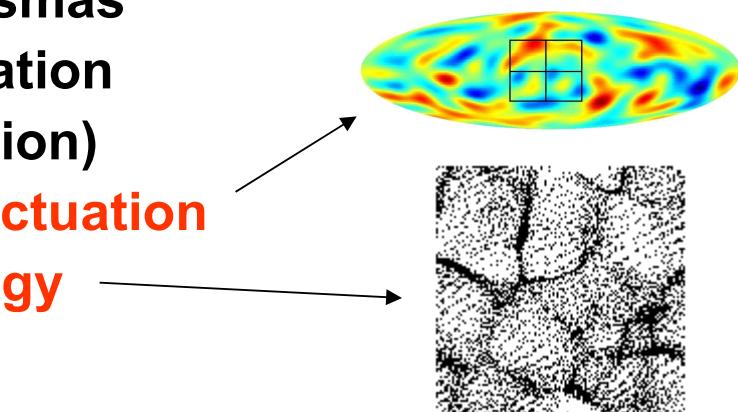
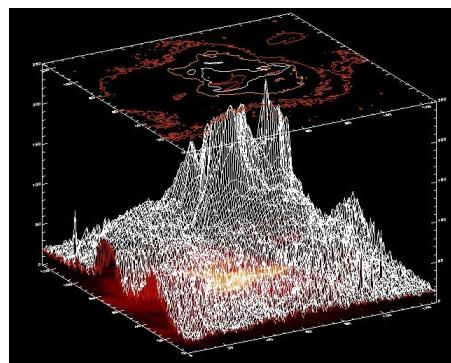
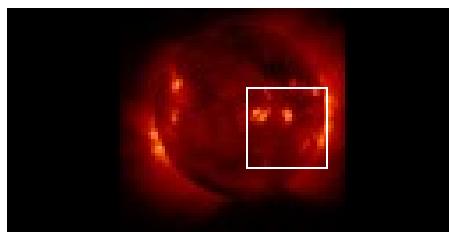
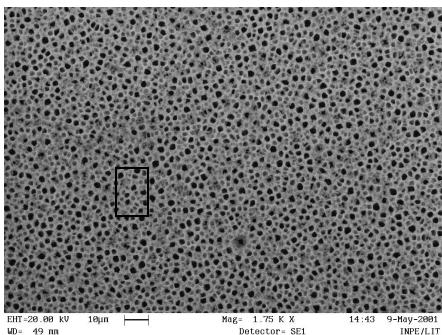
The resultant x of the equation represents the rate of rotation of the cylinder, y represents the difference in temperature at opposite sides of the cylinder, and the variable z represents the deviation of the system from a linear, vertical graphed line representing temperature.



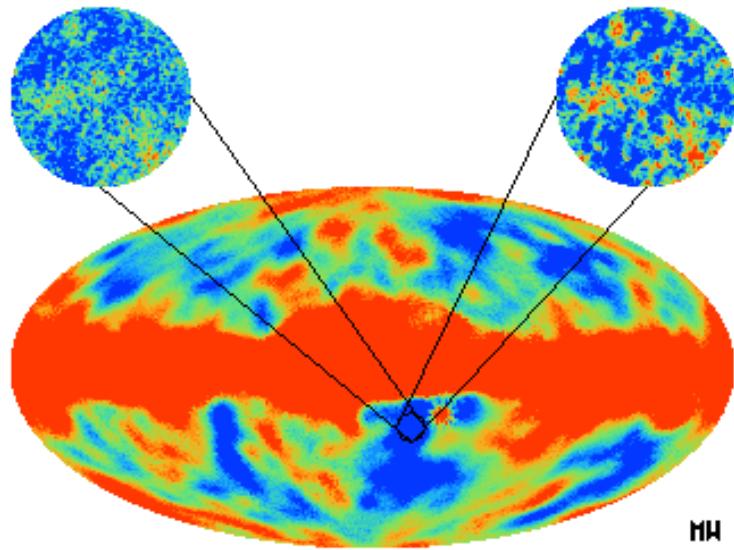


Computational Tools for Data Analysis:

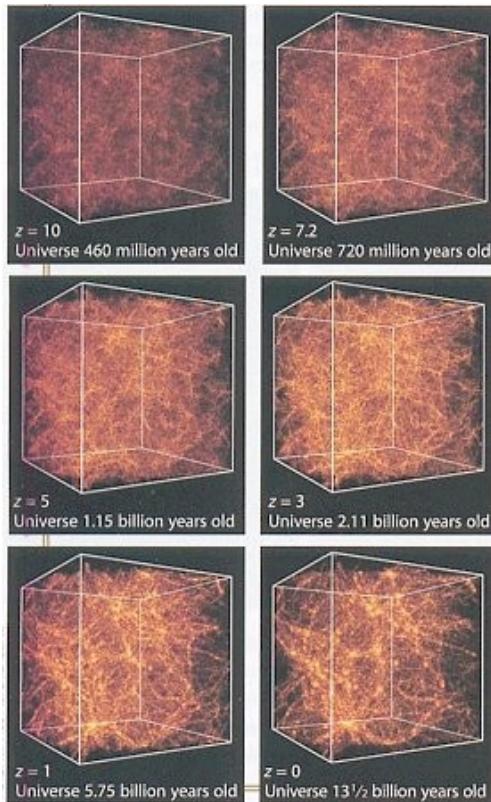
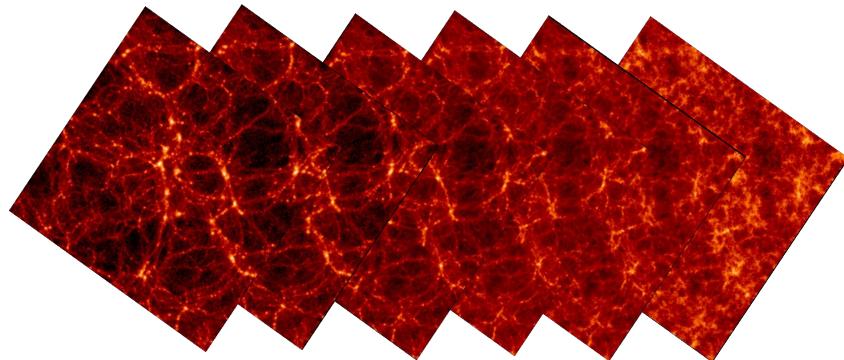
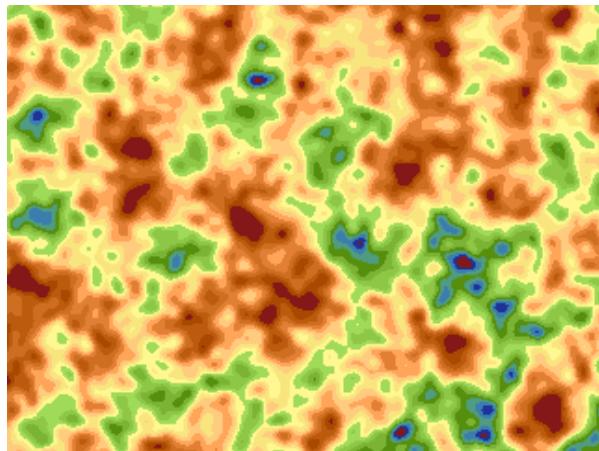
- Pattern characterization for nanofabrication
- Pattern evolution in astrophysical plasmas
- Pattern classification of wave propagation in elastic manifolds (membrane vibration)
- Pattern recognition of temperature fluctuation and large scale structures in cosmology



02 Projects: (1) Topological Signature ; (2) Turbulent-like model (poster)



HI



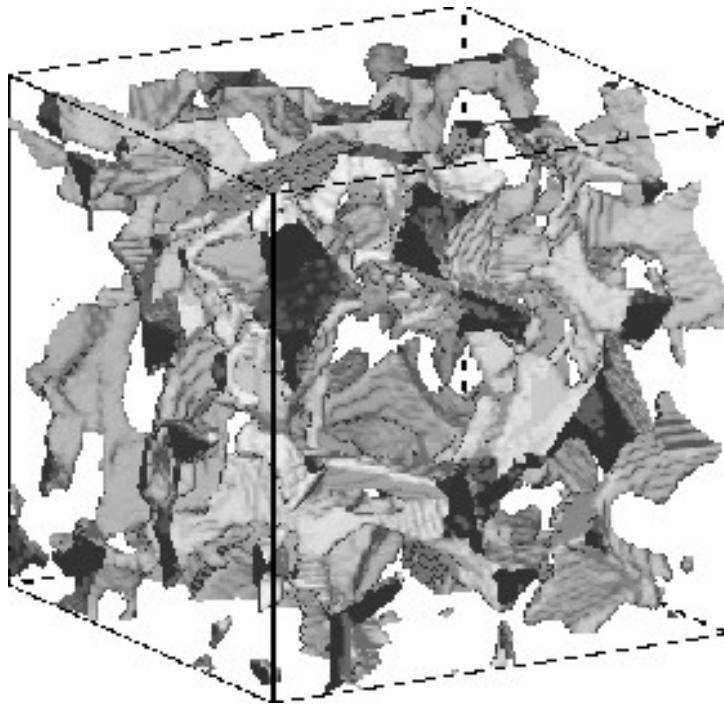
Analysing Regular and Irregular Patterns

- 1. Searching for Regularities:
2D FFT and 2D Wavelets Analysis**
- 2. Searching for Irregularites:Gradient
Pattern Analysis e MF**

Advanced Techniques (2 & 3D)

- Spatial Correlation Function and Statistical Exponents (universality classes)
- • Aspect Ratio and Minkowski Functionals (“integral geometry”)
- Geometric Dimensions
- Disorder Function
- • Gradient Pattern Analysis (“convex geometry”)
- Structural Entropy

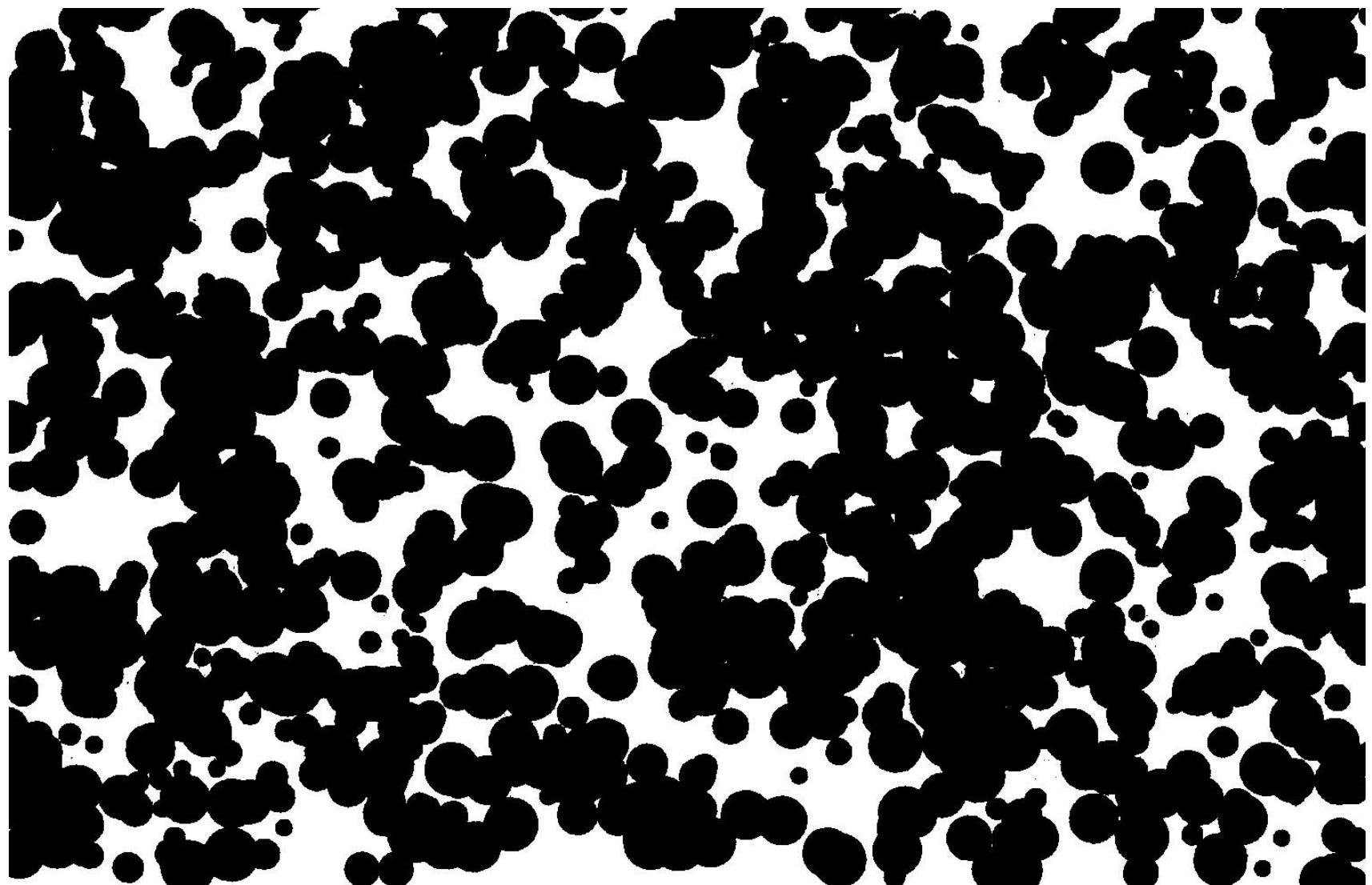
Integral Geometry in Physics: the shape of matter



3D-structure information in real space

Minkowski functionals

Boolean model: overlapping spheres



Minkowski functionals of black region ?

From Prof. Mecke

Gaussian random fields: Minkowski functions

Minkowski functionals : $f_v(\rho) \sim \lambda^v H_{v-1} [\rho(x)]$

with the Hermite functions $H_{v-1}(x) = (-\partial/\partial x)^{v-1} (1/2\pi) e^{-1/2x}$

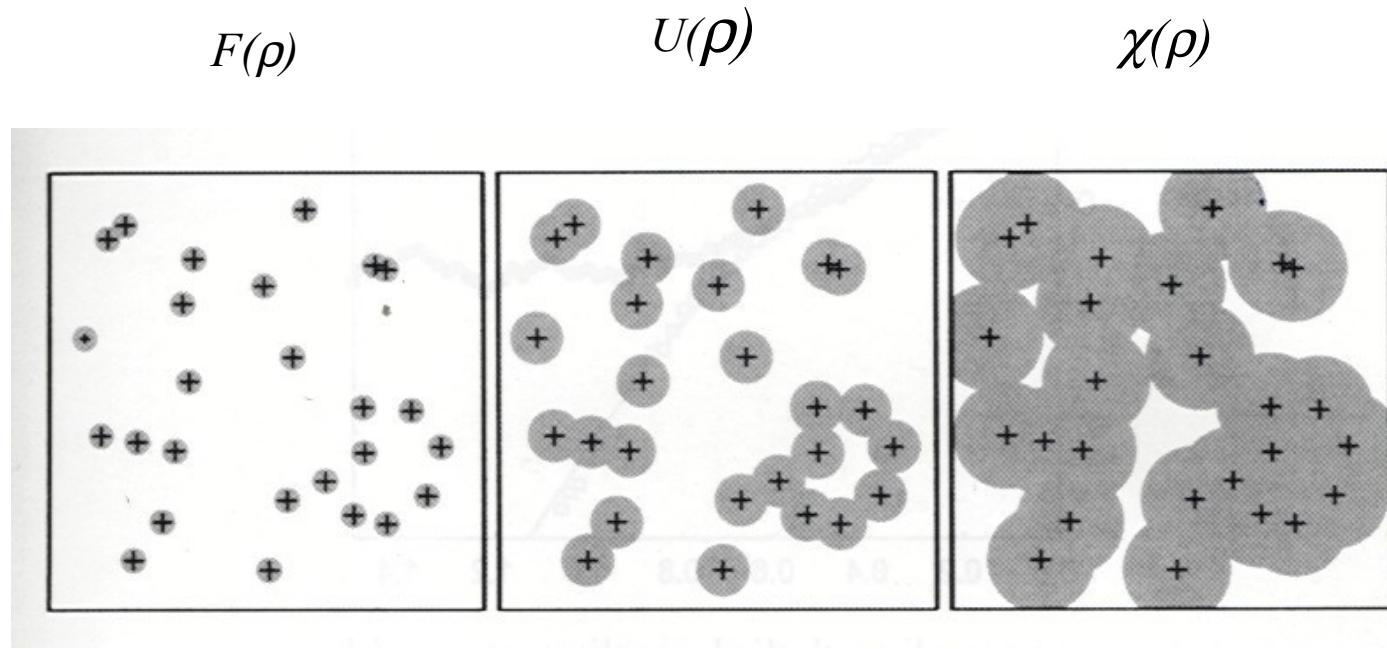
in two dimensions

area $F(\rho) = 1 - \frac{1}{\sqrt{2\pi}\sigma} \int_{-\infty}^{\rho} dx e^{-\frac{1}{2}(x-\rho_0)^2/\sigma^2}$

boundary $U(\rho) = \sqrt{\frac{\pi}{2}} \lambda e^{-\frac{1}{2}(\rho-\rho_0)^2/\sigma^2}$

curvature $\chi(\rho) = \frac{1}{\sqrt{2\pi}} \lambda^2 \frac{\rho-\rho_0}{\sigma} e^{-\frac{1}{2}(\rho-\rho_0)^2/\sigma^2}$

Surface (2D) into Volume (3D):

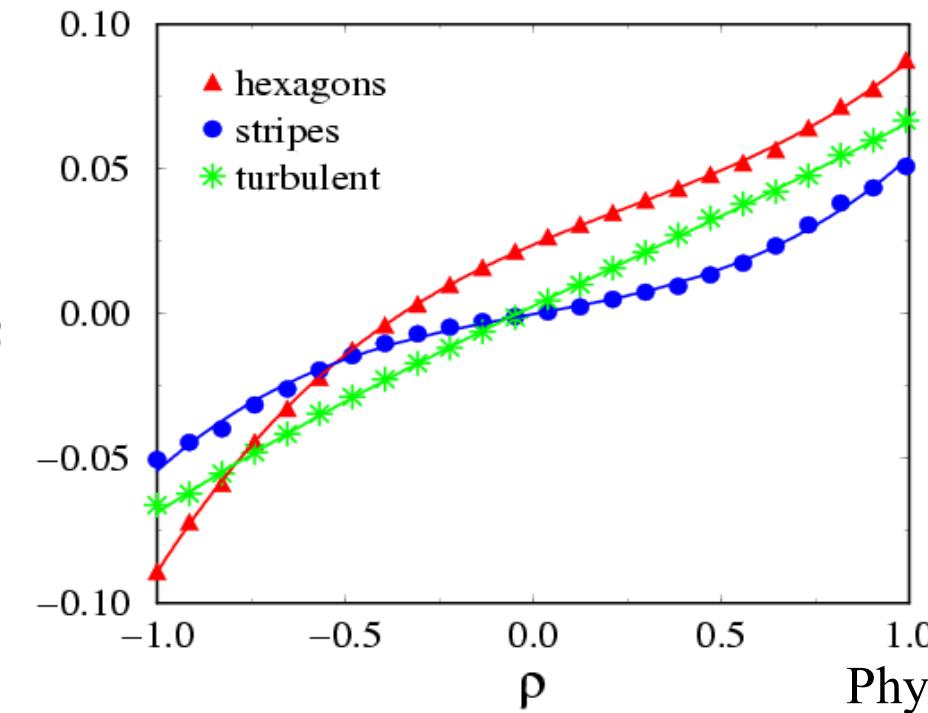
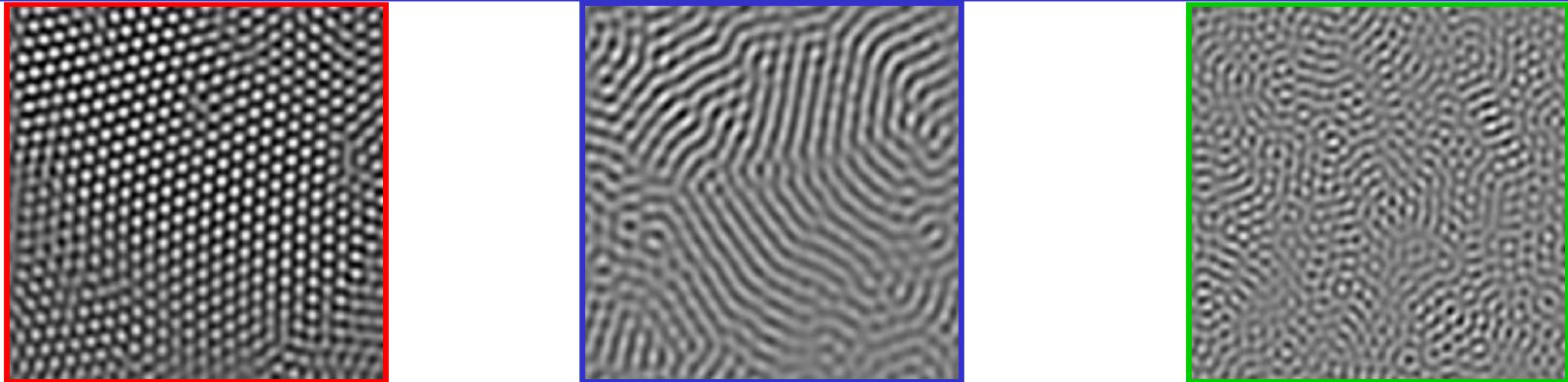


$F(\rho)$ = superficie de cobertura

$U(\rho)$ = contorno

$\chi(\rho)$ = característica de Euler
(conectividade)

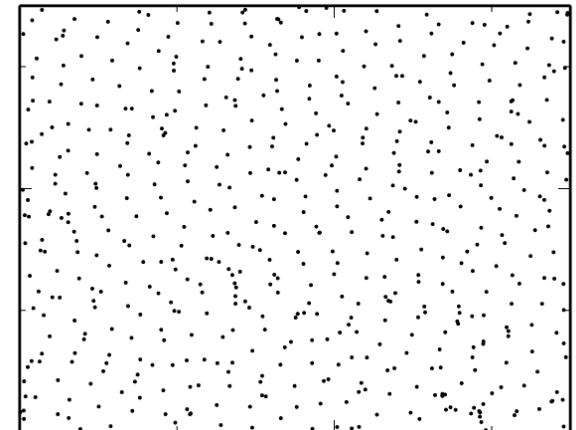
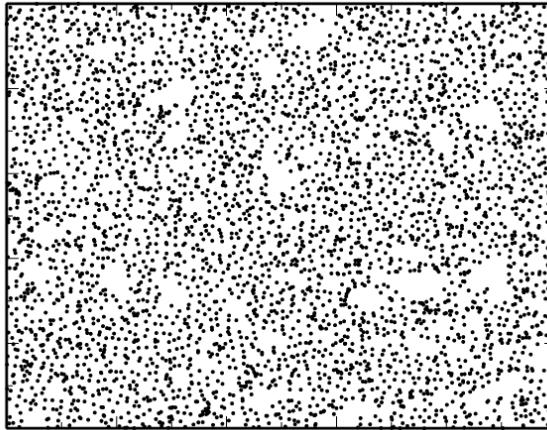
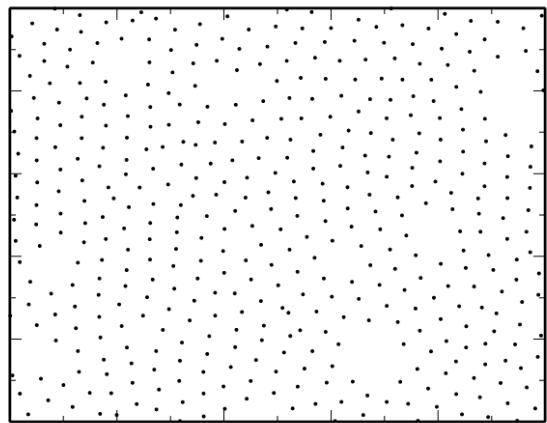
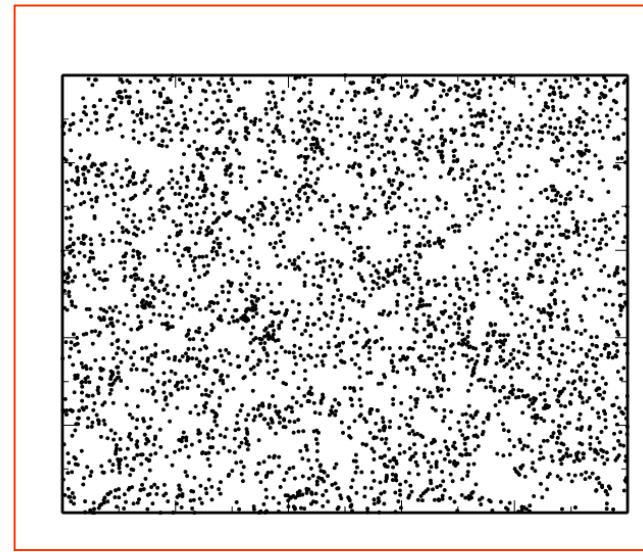
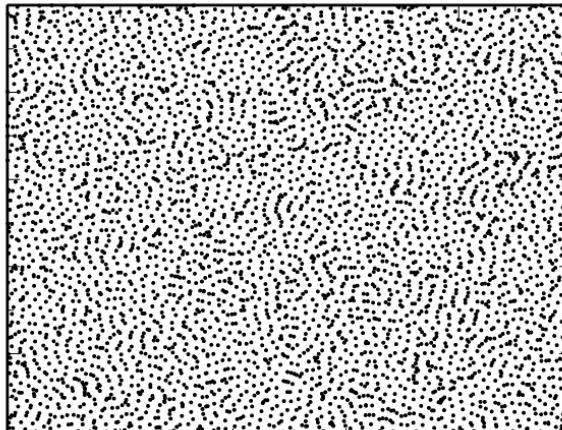
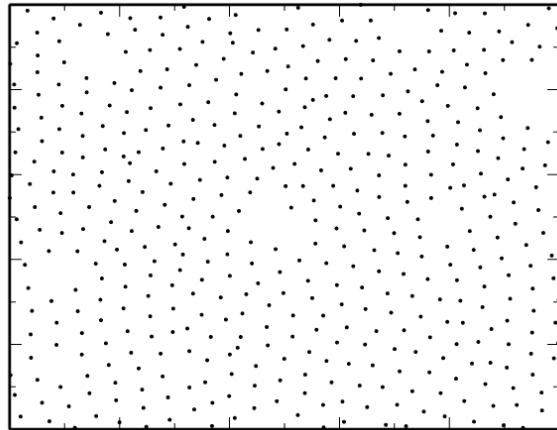
Turing patterns: Euler characteristic



Phys. Rev. E 53,4794 (1996)

Minkowski functions are cubic polynomials in threshold ρ (From Prof. Mecke)

At this moment: simulating complex fluid configurations



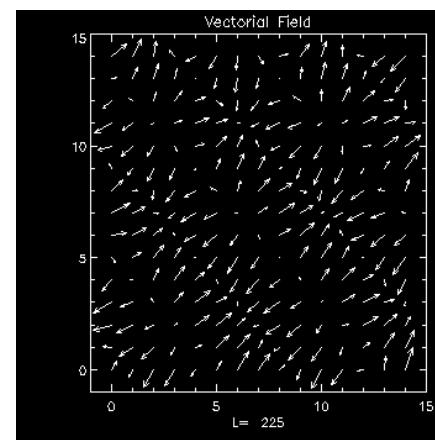
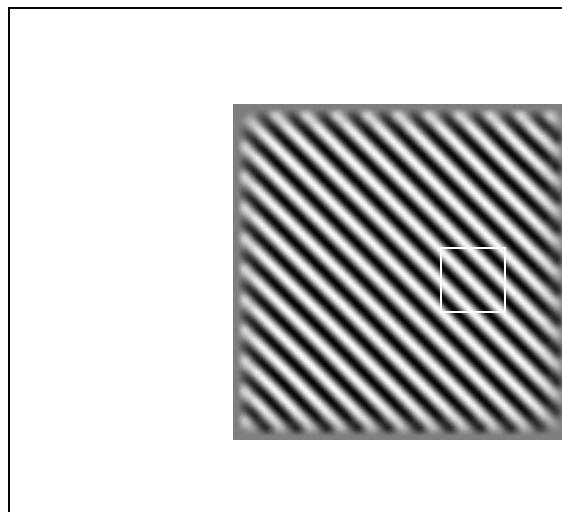
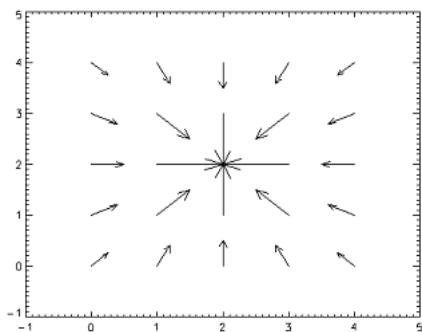
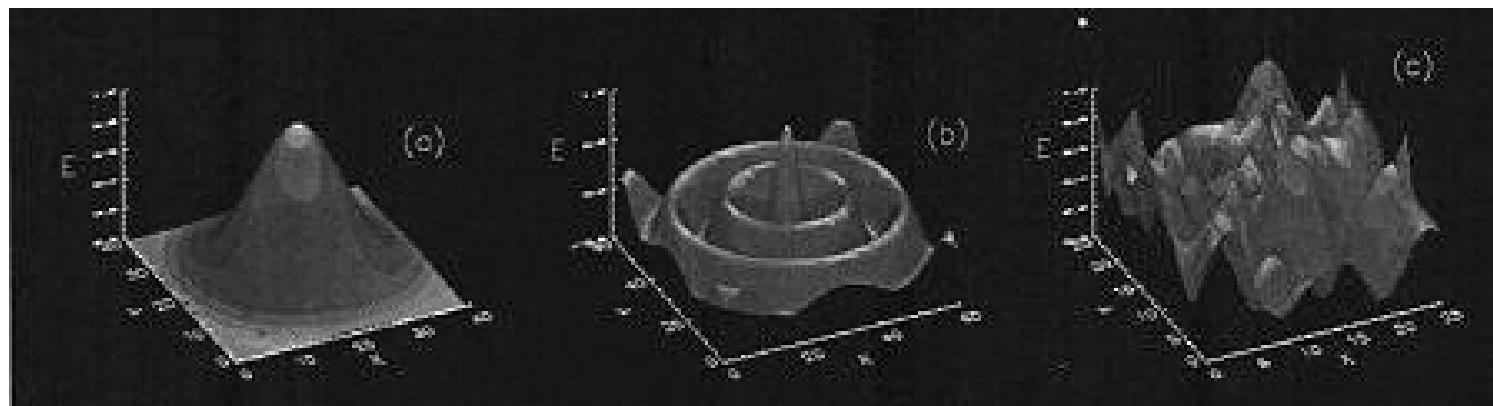
What about asymmetries?

From Prof. Mecke

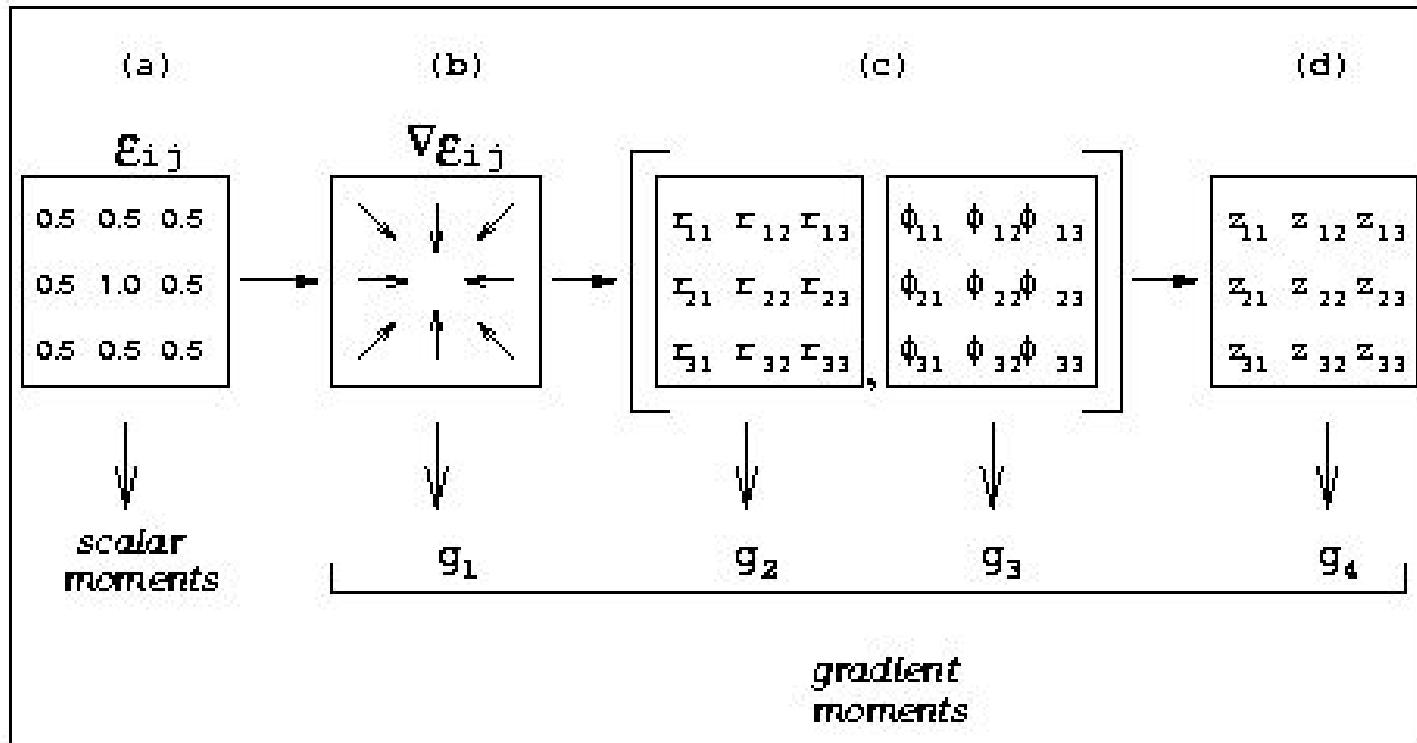
The Notion of Local & Global Fluctuations

4.56	4.67	4.59	5.01	5.03	6.98	6.98	6.87	5.67	6.51	5.00
4.00	4.04	4.09	4.34	4.56	4.34	4.52	5.56	5.78	4.56	5.67
3.88	3.45	3.56	6.45	6.34	6.89	5.09	5.56	5.77	5.67	5.68
3.34	3.56	4.54	4.00	4.04	5.08	5.20	6.34	5.34	4.56	4.59
5.56	4.56	4.59	5.11	5.23	5.67	5.23	6.02	8.79	8.55	9.00
7.89	8.09	7.29	8.08	7.34	6.78	6.79	9.08	0.08	0.06	0.89
9.02	9.08	9.08	9.04	9.01	8.79	8.88	8.88	0.09	0.06	1.24
9.08	9.06	9.78	9.89	9.56	9.05	9.08	9.06	1.23	1.34	0.09
4.56	4.67	4.59	5.01	5.03	6.98	6.98	6.87	1.22	1.45	0.09
4.00	4.04	4.09	4.34	4.56	4.34	4.52	5.56	1.34	1.56	1.01
3.68	3.45	3.16	6.45	5.34	6.19	5.09	5.56	1.56	1.56	1.55
5.14	3.56	3.14	6.00	4.04	5.08	7.22	9.34	4.56	4.01	4.45
5.56	4.56	4.59	2.11	5.23	5.67	9.23	9.00	5.67	5.01	6.09
5.98	4.03	3.09	3.01	7.32	6.78	5.69	9.01	8.00	7.78	6.00

Symmetry Breaking in the Gradient Field of the Amplitudes → GPA



Gradient Pattern Analysis



$\mathbf{G}_t = \nabla[E(x,y,t)_t]$ is represented by 4 n.s. **gradients moments**:

$$g_1^{(t)} \equiv h_1 (\{(r_{1,1}, \phi_{11}), \dots, (r_{ij}, \phi_{ij}), \dots, (r_{kk}, \phi_{kk})\}_t)$$

$$g_2^{(t)} \equiv h_2 (\{(r_{11}), \dots, (r_{ij}), \dots, (r_{kk})\}_t)$$

$$g_3^{(t)} \equiv h_3 (\{(\phi_{11}), \dots, (\phi_{ij}), \dots, (\phi_{kk})\}_t)$$

$$g_4^{(t)} \equiv h_4 (\{(z_{11}), \dots, (z_{ij}), \dots, (z_{kk})\}_t)$$

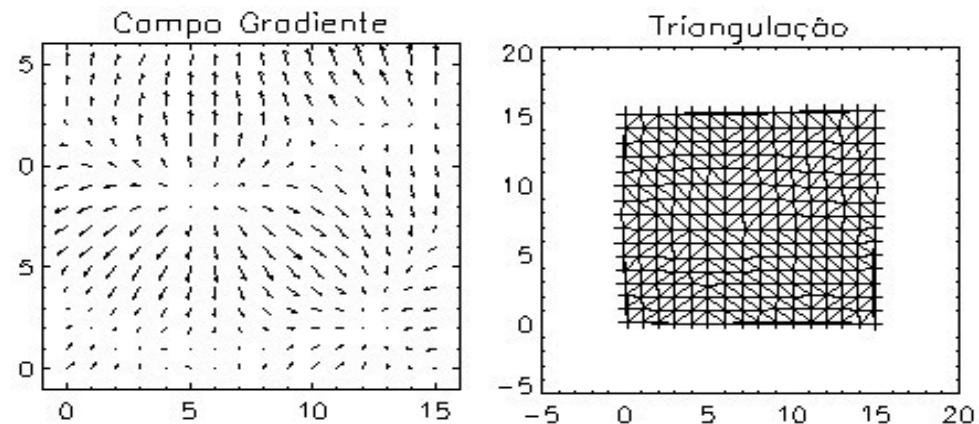
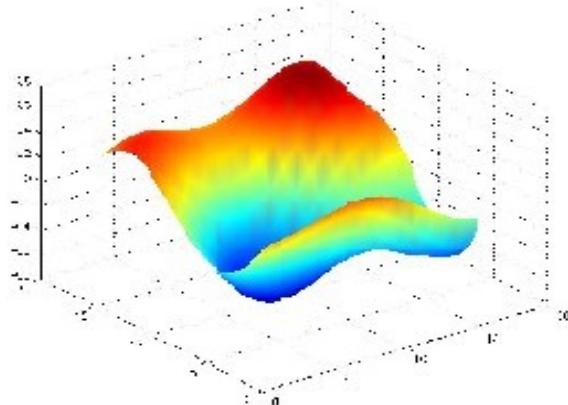
As $\{r\}$ and $\{\phi\}$ are compact groups, spatially distributed, they can be geometrically constructed as *Haar-like* measures ==> rotational and amplitude translation invariant

How to compute g_1, g_2, g_3 e g_4 ?

$$g_1 \equiv (\mathbf{C} - \mathbf{V}_A)/\mathbf{V}_A \quad , \quad \mathbf{C} > \mathbf{V}_A$$

V_A = amount of asymmetric vectors, ($v_1 + v_2 \neq 0$)

C = amount of geometric correlation lines (Delaunay triangulation $T_D(C, V_A)$)
*(Asymmetric Amplitude Fragmentation - AAF, Rosa et al.,
Int. J. Mod. Phys. C, 10(1)(1999):147.)*



$$g_2 \equiv \sum (r_{ij} - r_{mn})^2 / N ;$$

$$g_3 \equiv \sum (\Phi_{ij} - \Phi_{mn})^2 / N$$

$g_4:$ $S_z = - \sum z_{i,j}/z \ln (z_{i,j}/z) =$

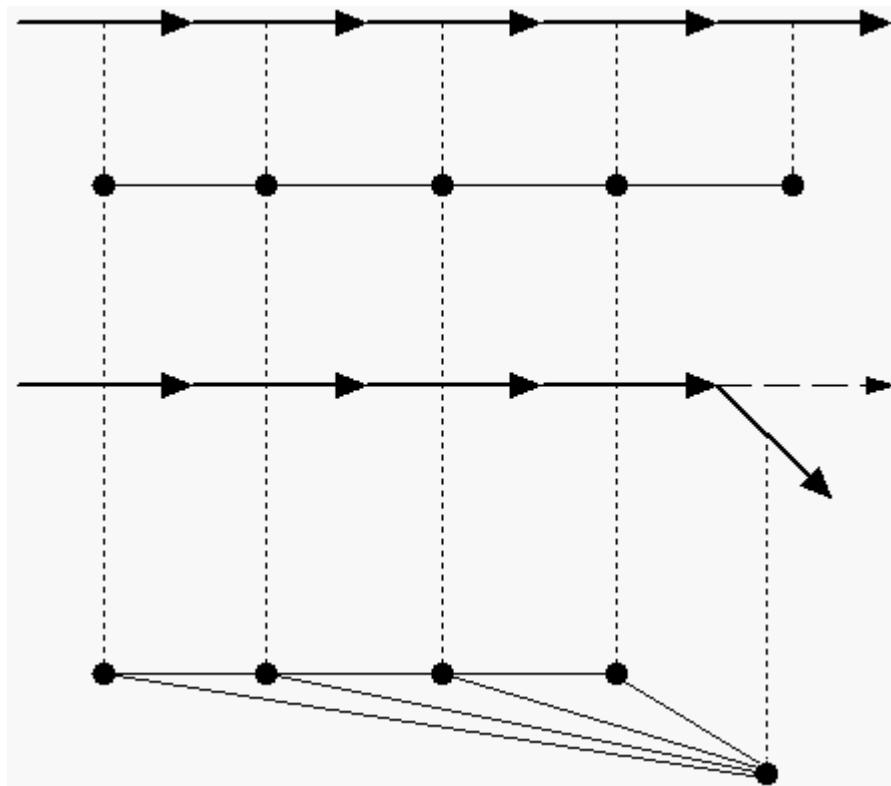
$$\text{Re}(S_z) e^{i S_z}$$

$|g_4| = \text{Re}(S_z) = S(|z|) \text{ and } \Phi(g_4) = \text{Im}(S_z)$

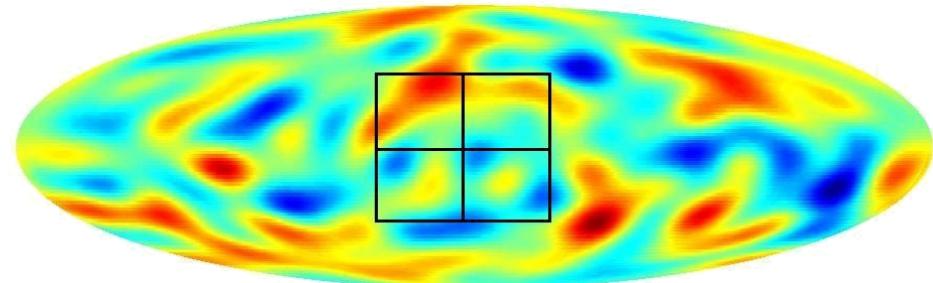
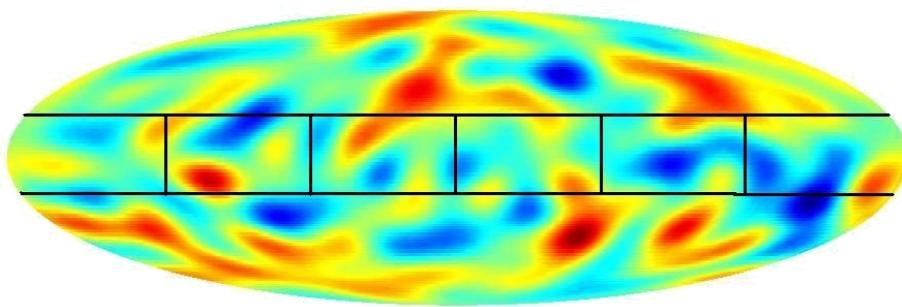
Thus, $|g_4|$ and $\Phi(g_4)$ are invariant measurements of norm and phase of the gradient entropy

(*Complex Entropic Form* (CEF) by Ramos et al. Physica A283(2000):171.)

GPA: Rosa et al. Braz. J. of Physics, 2003

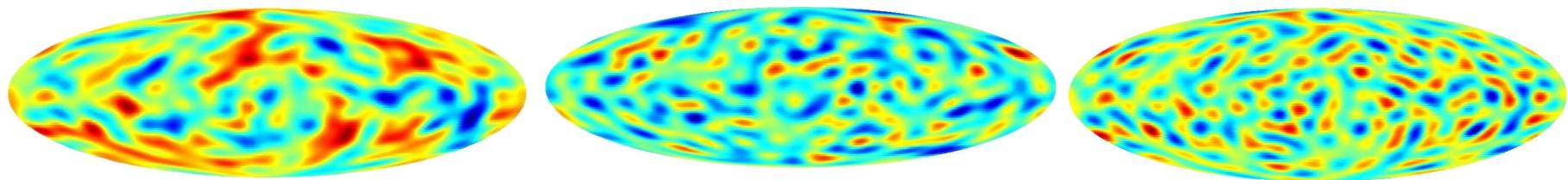


GPA of Temperature Fluctuations on a Hariisson-Zeldovich Sphere



Modelling Irregular Amplitude Fluctuations on a Spherical Surface

$$\frac{\delta A}{A}(\theta, \varphi) = \frac{A(\theta, \varphi) - A_0}{A_0}$$

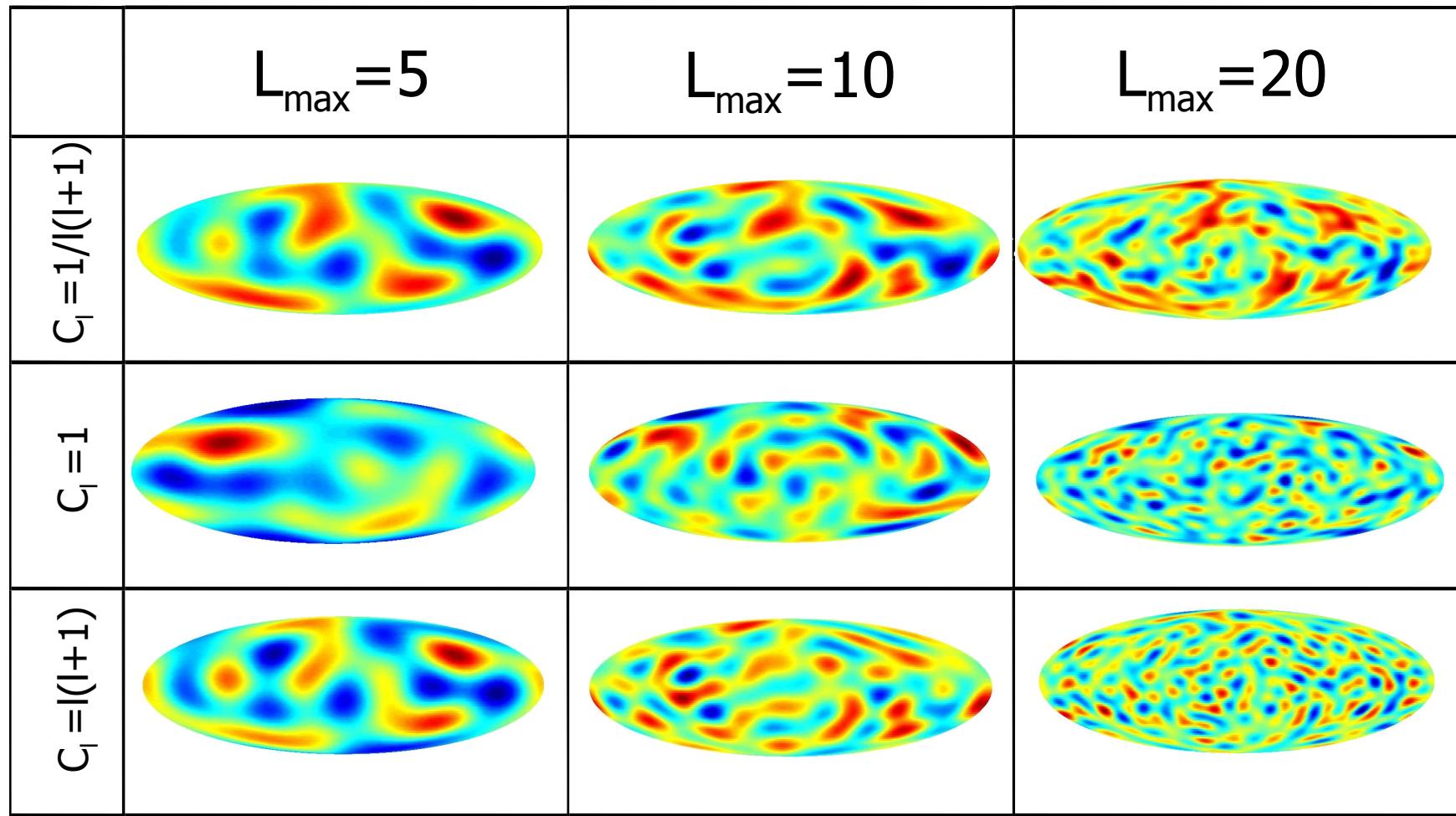


$$\frac{\delta T}{T}(\theta, \varphi) = \sum_{l=0}^{\infty} \sum_{m=-l}^l a_{lm} Y_{lm}(\theta, \varphi)$$

- ◆ Plane wave decomposition in spherical harmonics:

$$e^{i\vec{k} \cdot \vec{r}} = 4\pi \sum_{lm} i^l jl(kr) Y_{lm}^*(\theta_{\vec{k}}, \varphi_{\vec{k}}) Y_{lm}(\theta, \varphi)$$

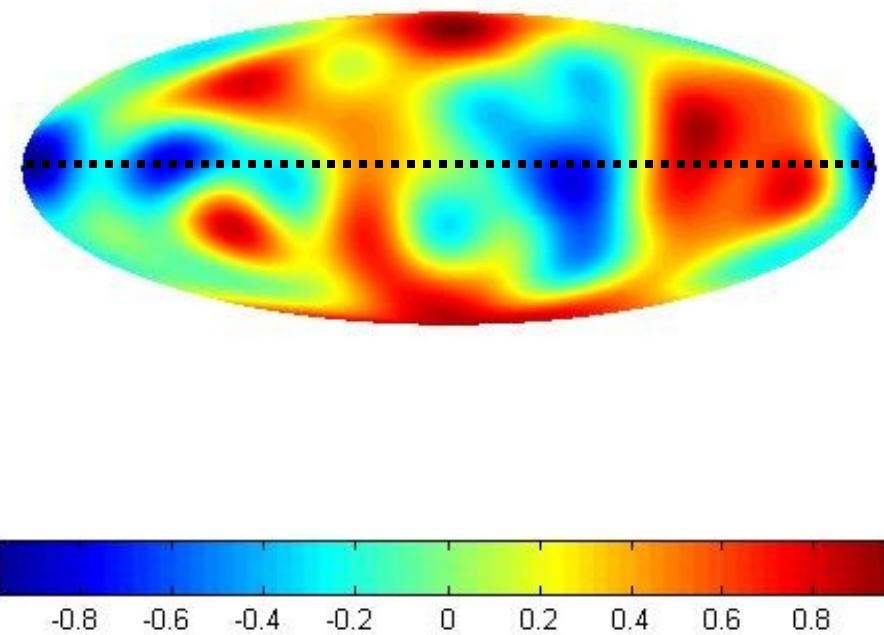
Irregular Fluctuations:



Asymmetric Aspect Ratio (AAR)

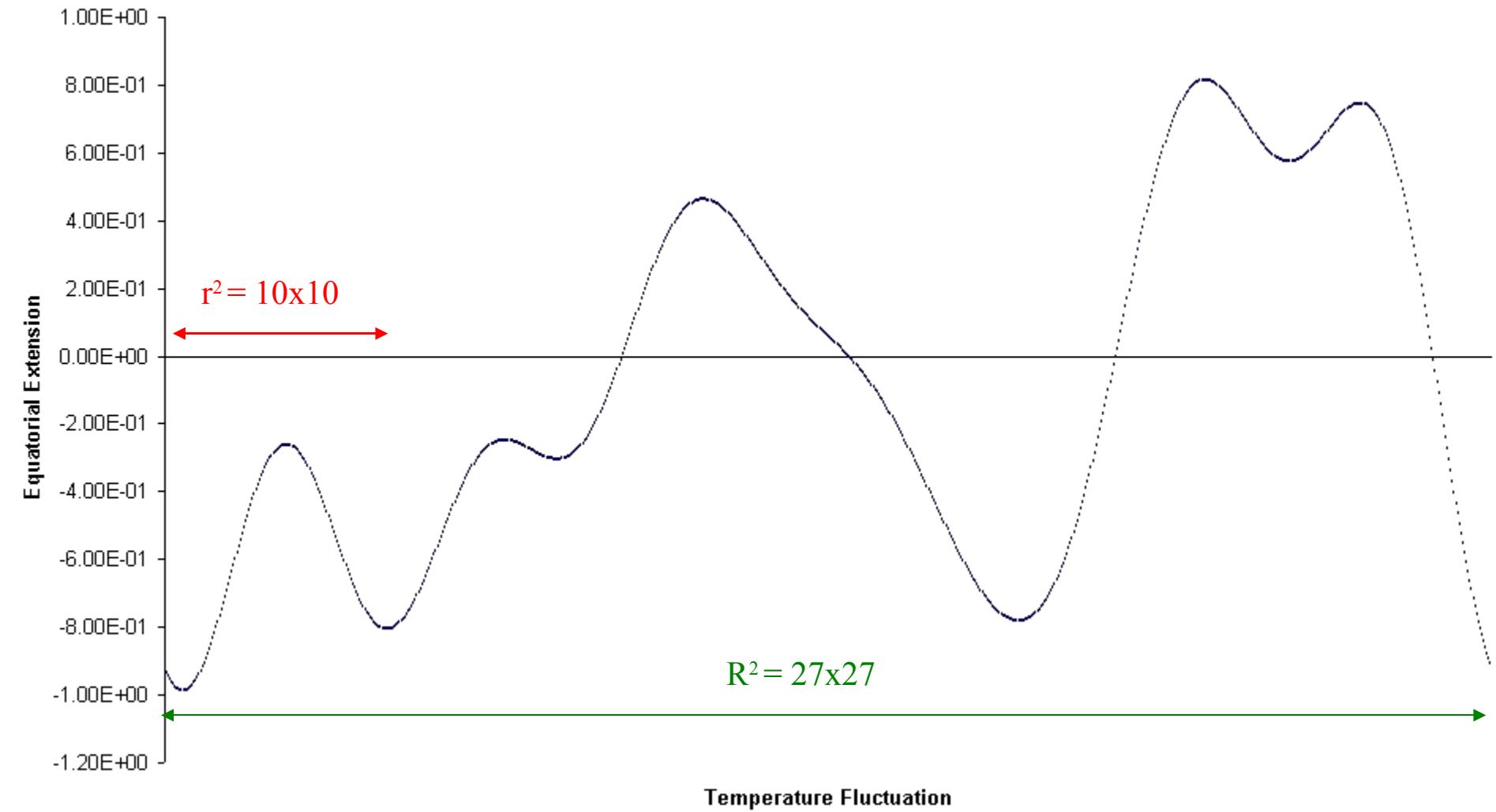
- In our approach we take from each CMB map a relation between the maximum variation of asymmetries for two characteristic scales: **the global scale R** where the major matrix ($R \times R$) is obtained and **the first local scale r** for that the values of g_a are maximally different when different random seeds are considered in the simulation.
- Therefore, when we compare two different *Asymmetric Aspect Ratio* (AAR) where each one is coming from a different topology (usually, a nontrivial one -NT- in comparison to the trivial standard flat topology-T), we can conjecture the existence of a parameter **γ** whose values, when different than unity, can distinguish (or classify) different classes of nontrivial topologies:

$$\frac{\Delta g_{T,R}^a}{\Delta g_{T,r}^a} \approx \gamma \left(\frac{\Delta g_{NT,R}^a}{\Delta g_{NT,r}^a} \right)$$

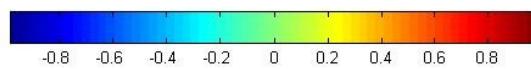
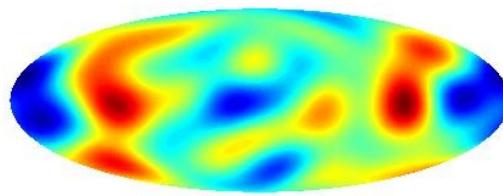
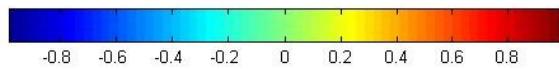
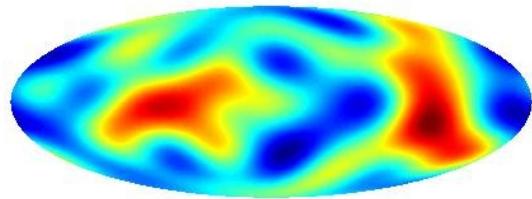
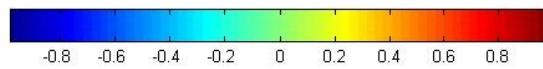
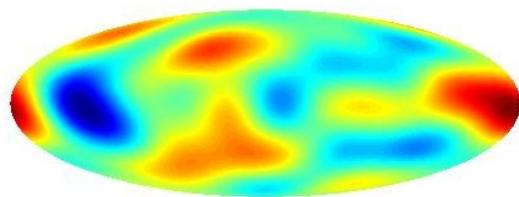
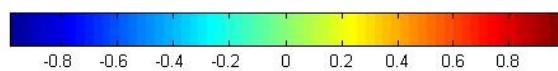
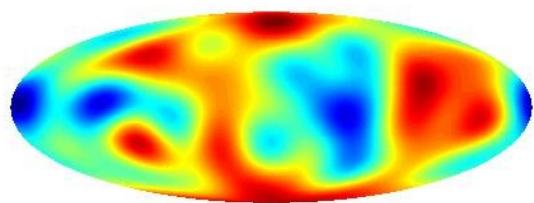


The equatorial fluctuation pixel profiles, at $f_\theta = 0$, as a function of the longitude sampled in 768 pixels ($R^2 = 27 \times 27$) (CT1)

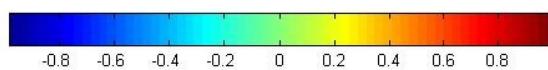
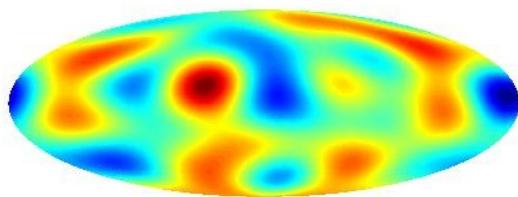
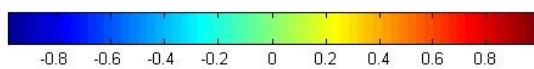
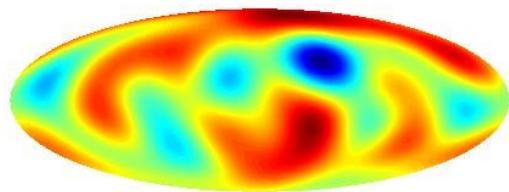
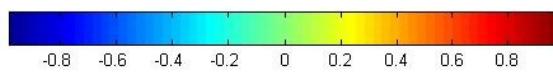
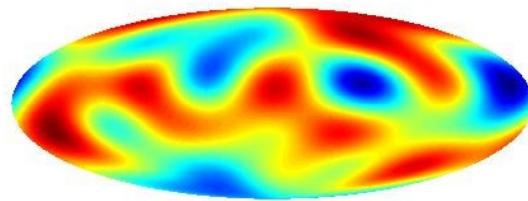
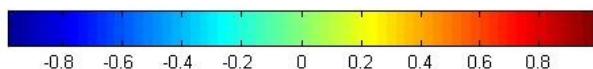
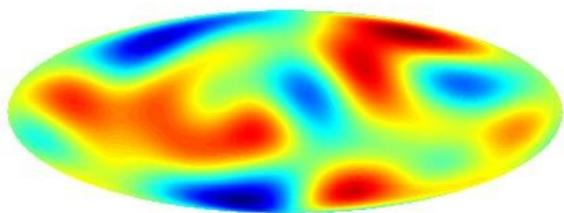
Non-Trivial Topology

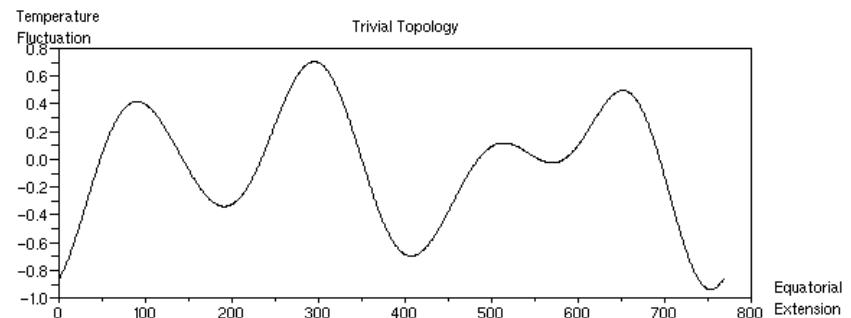
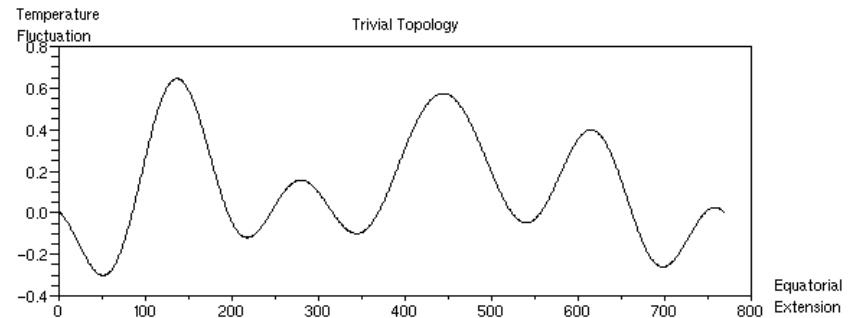
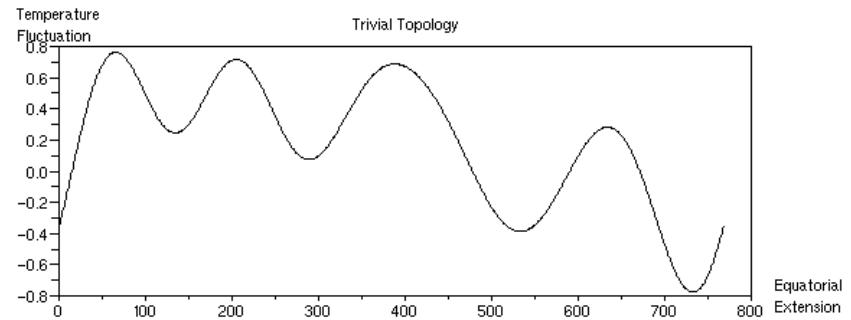
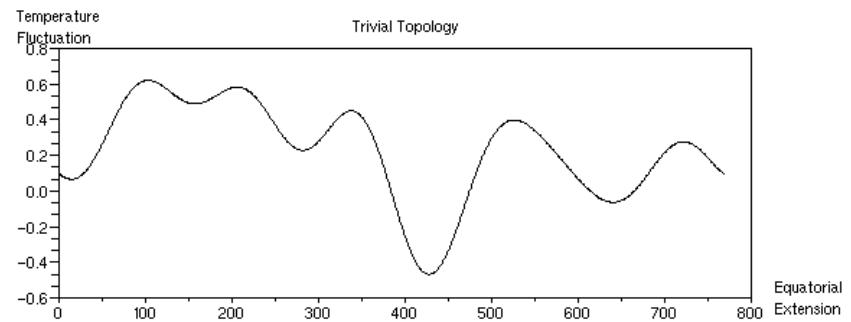
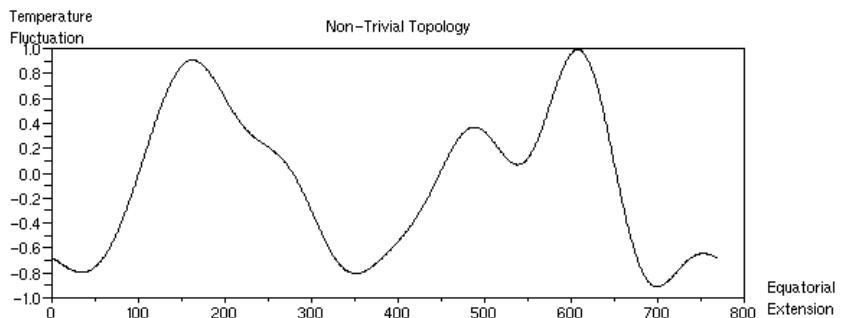
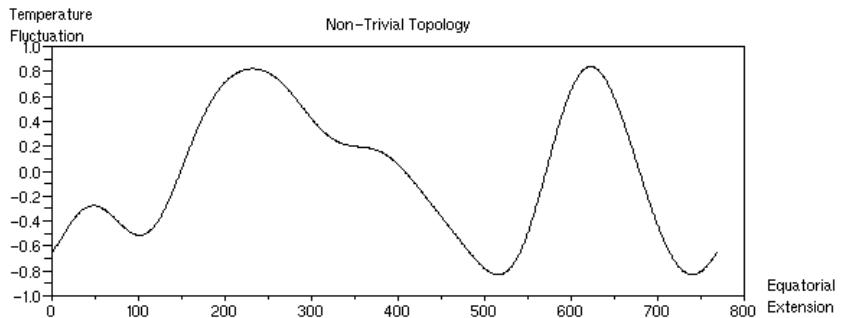
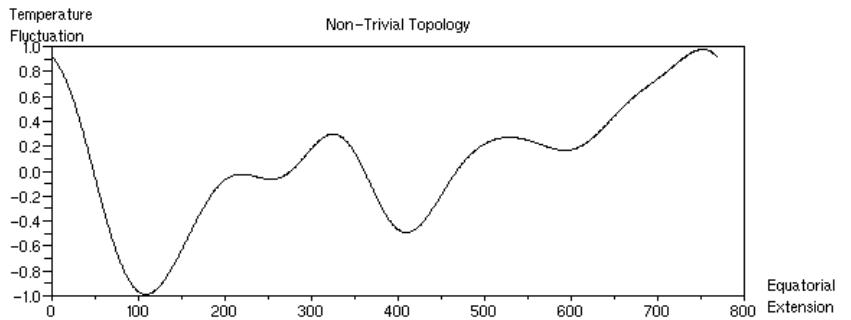
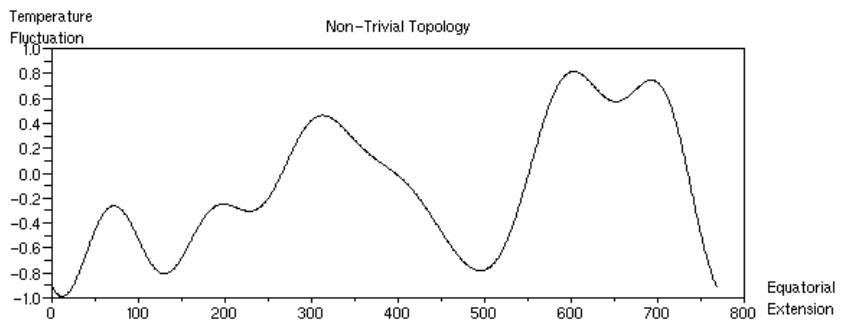


TC



TT



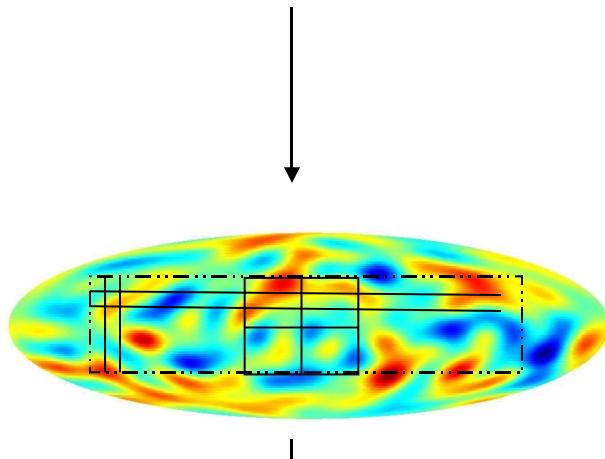


Resultados e Conclusões Preliminares

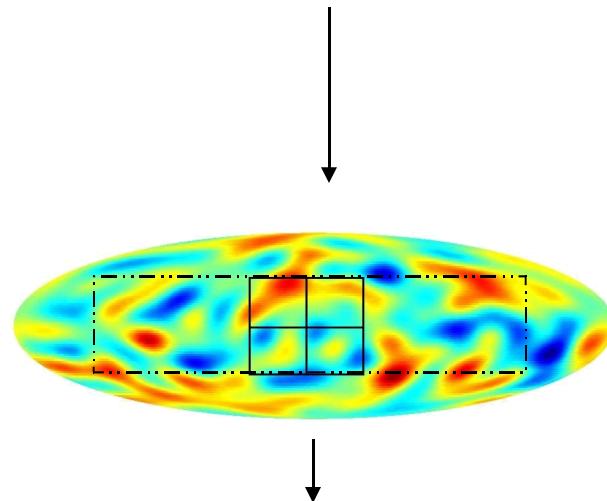
TC		TT	
$g_a(729)$	$g_a(100)$	$g_a(729)$	$g_a(100)$
1.96	1.79	1.94	1.80
1.96	1.81	1.95	1.81
1.98	1.87	1.92	1.80
1.95	1.74	1.99	1.87

$$\frac{\Delta g_{C,27}^a}{\Delta g_{C,10}^a} \approx 4 \left(\frac{\Delta g_{T,27}^a}{\Delta g_{T,10}^a} \right)$$

Trivial



Torus



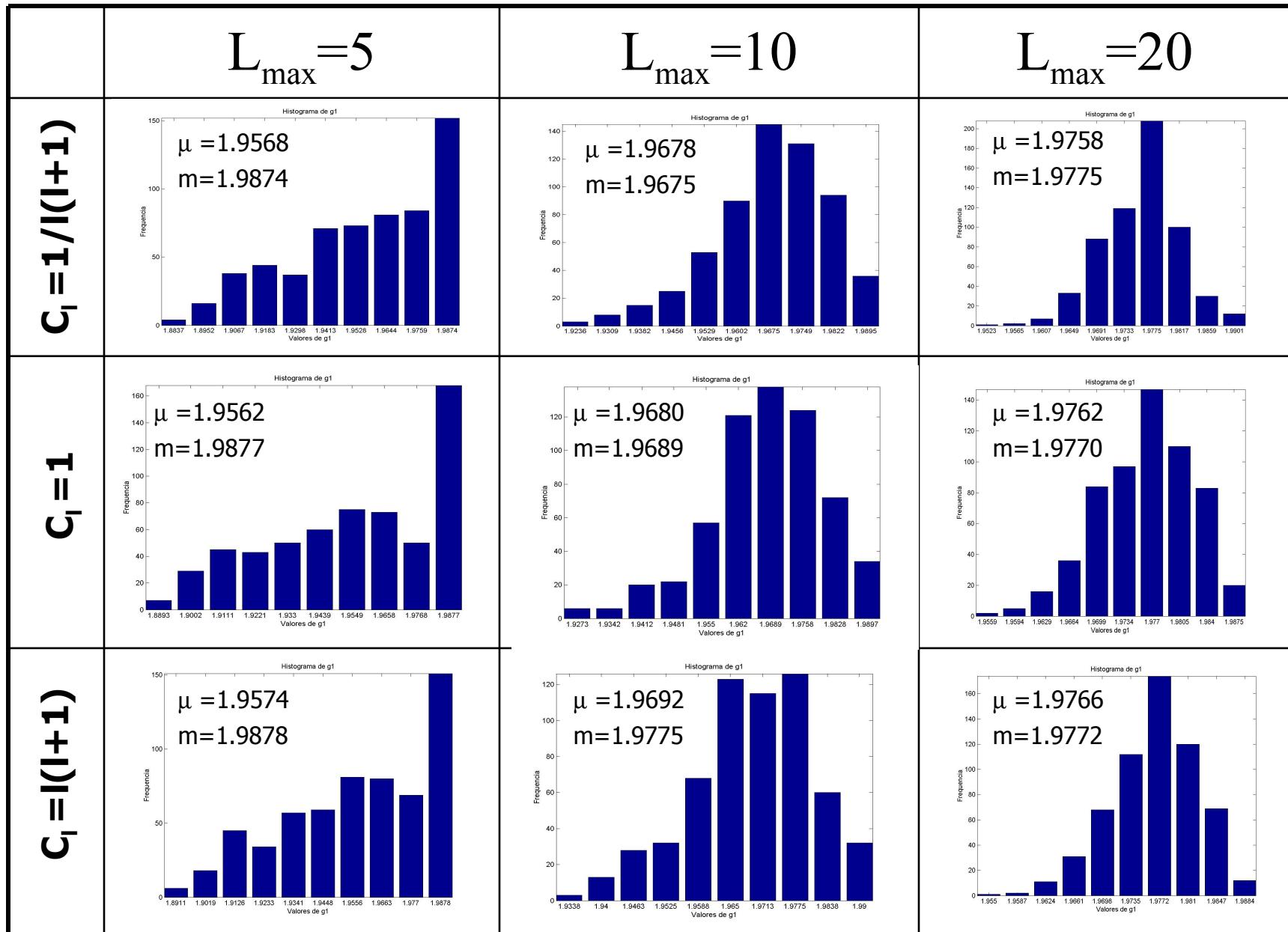
Gradient Classes $C_n(\Lambda\text{CDM})$

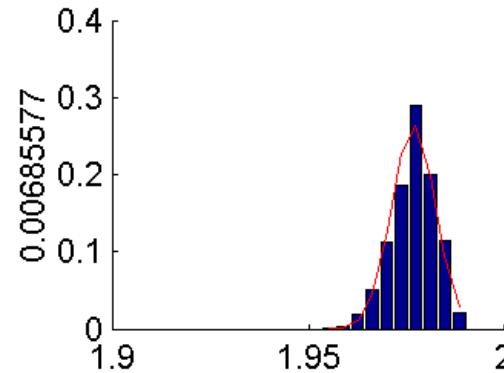
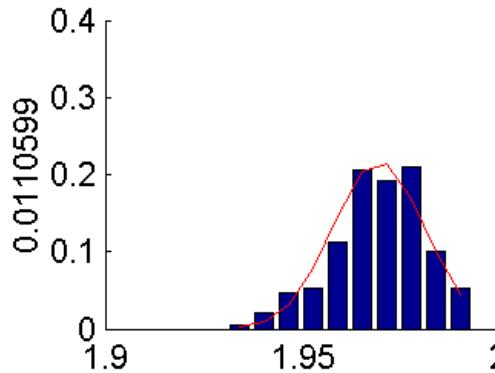
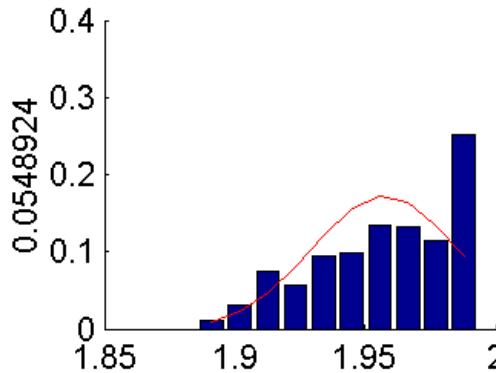
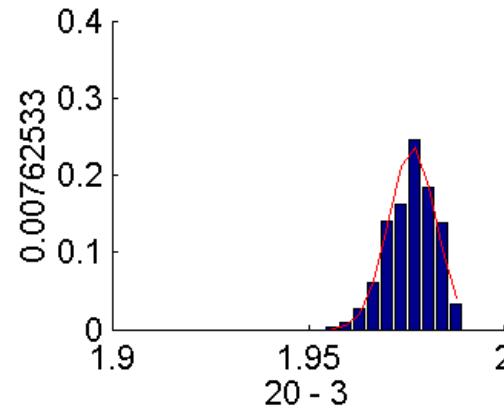
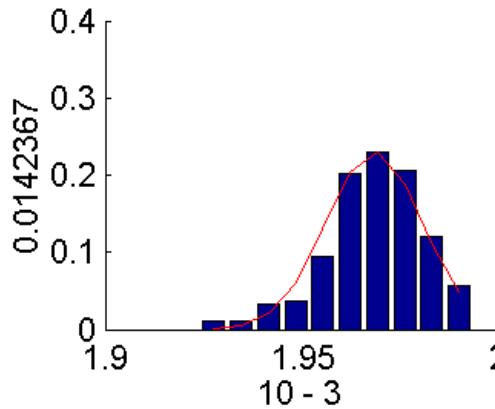
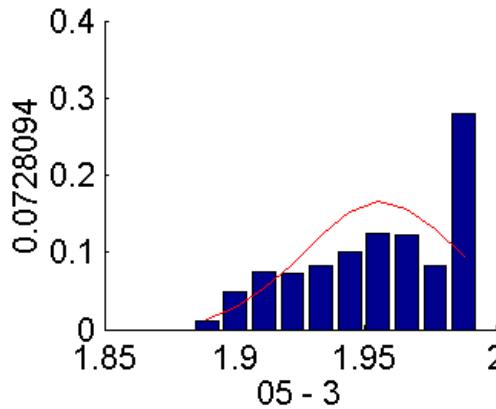
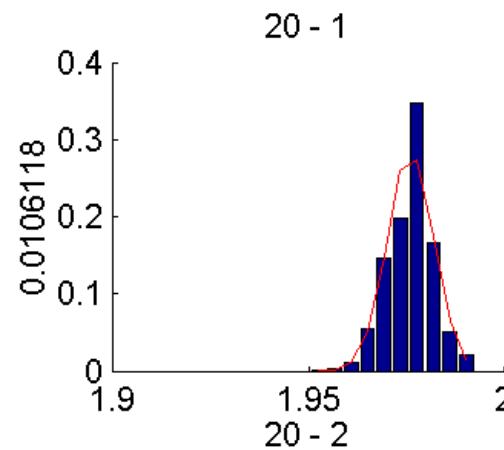
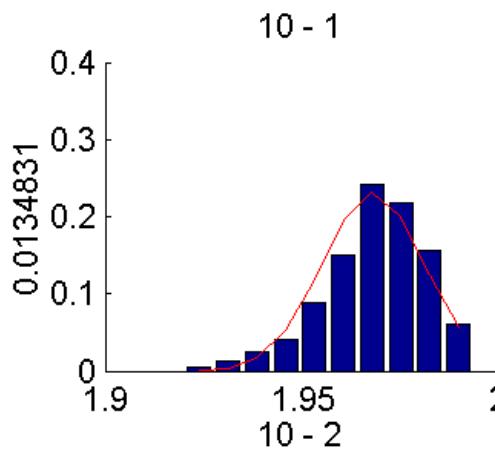
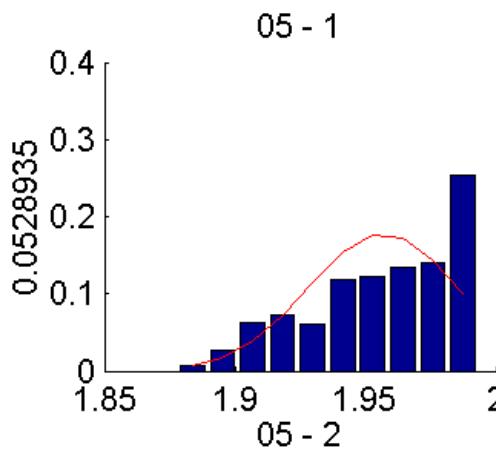
Gradient Classes $C_n(\text{Torus})$

\approx
?

REAL DATA ==> Gradient Classes C_n (ex. WMAP)

Results from the GPA of 900 simulations:

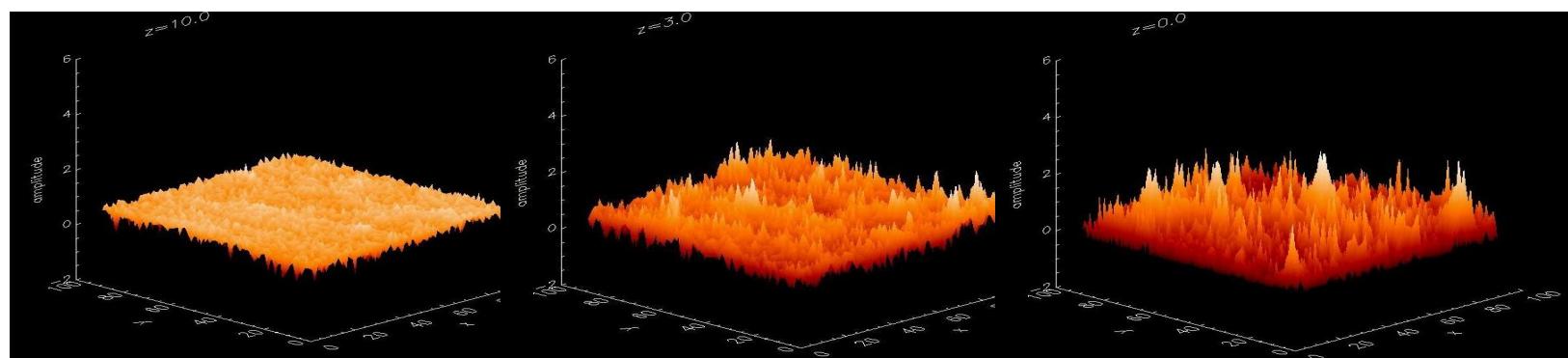
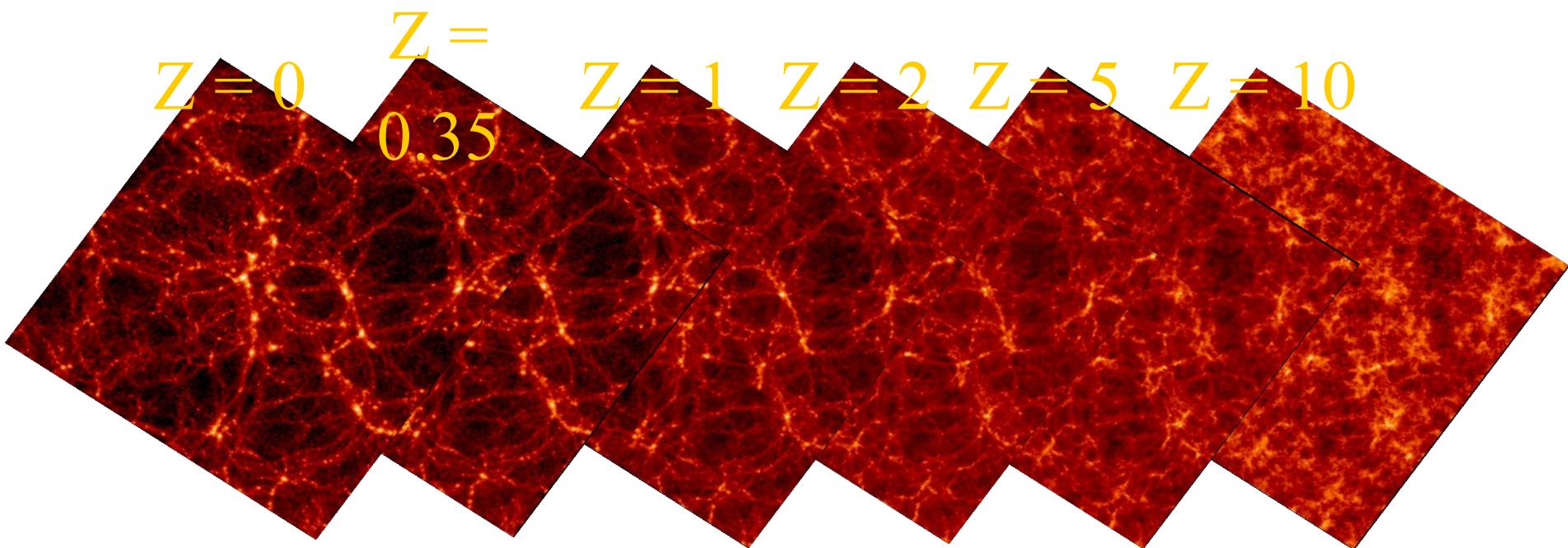


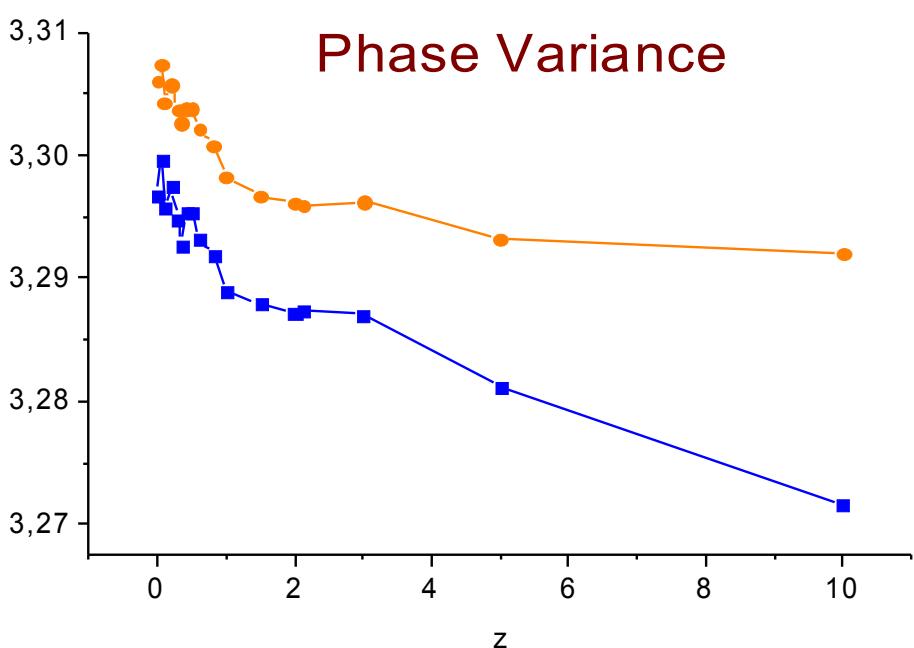
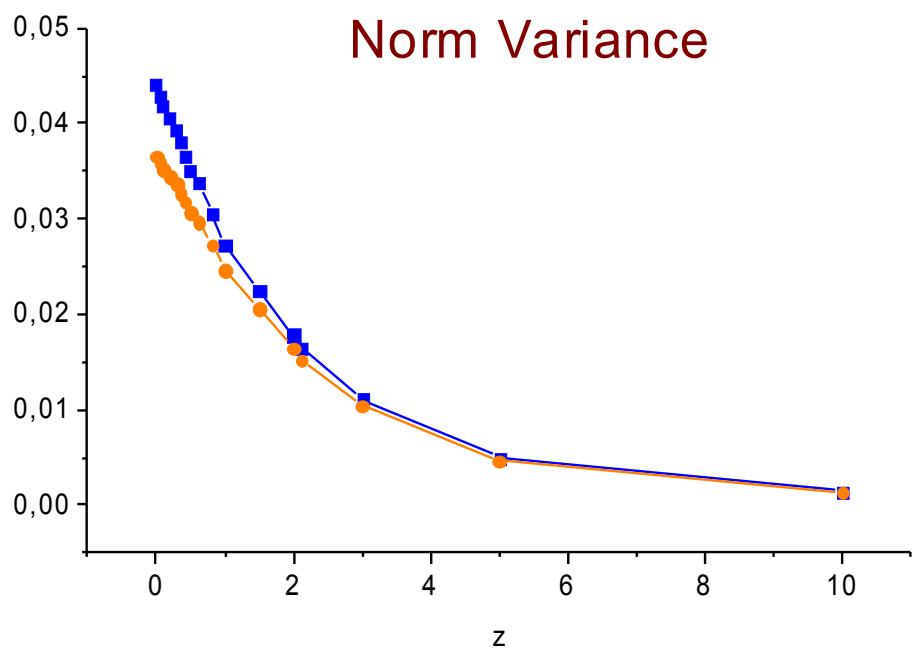
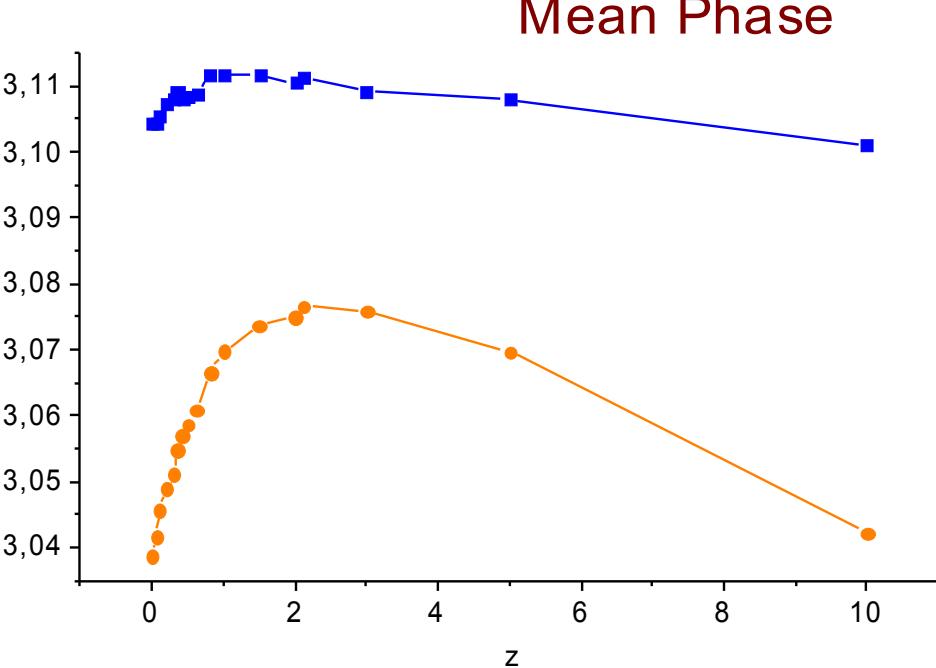
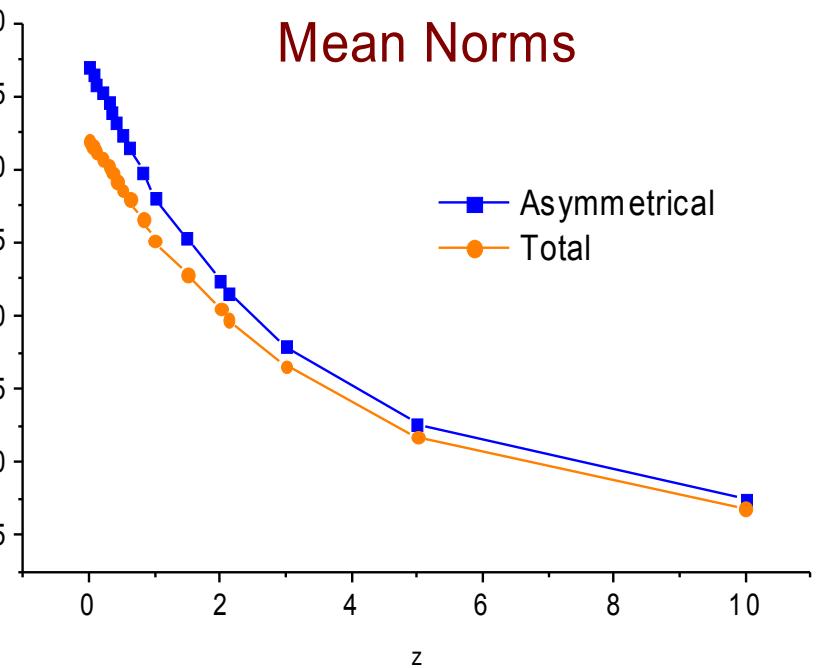


Resultados

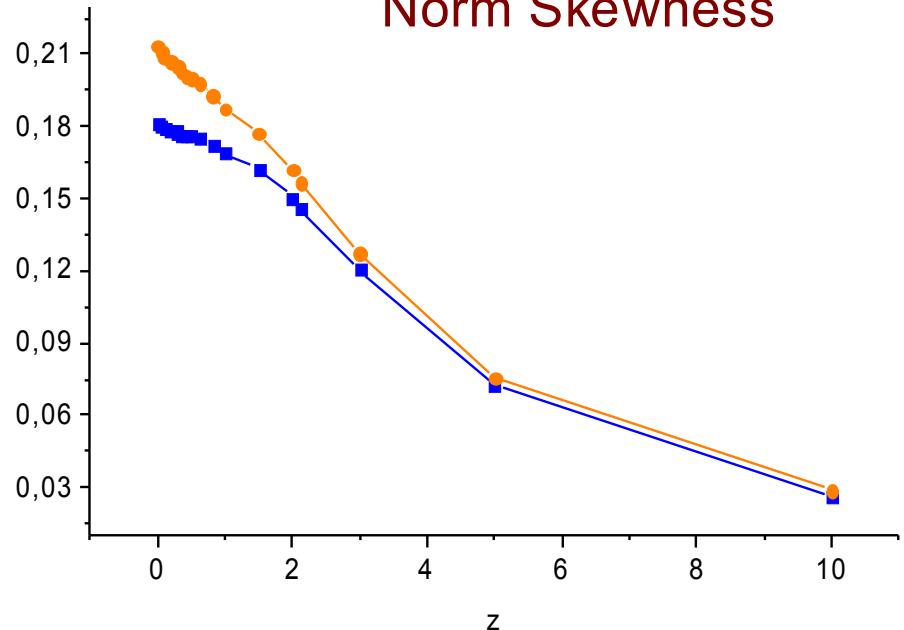
	Lmax=5	Lmax=10	Lmax=20
C1	0.053	0.014	0.011
C2	0.073	0.014	0.008
C3	0.056	0.011	0.007

GRADIENT PATTERN ANALYSIS IN STRUCTURE FORMATION

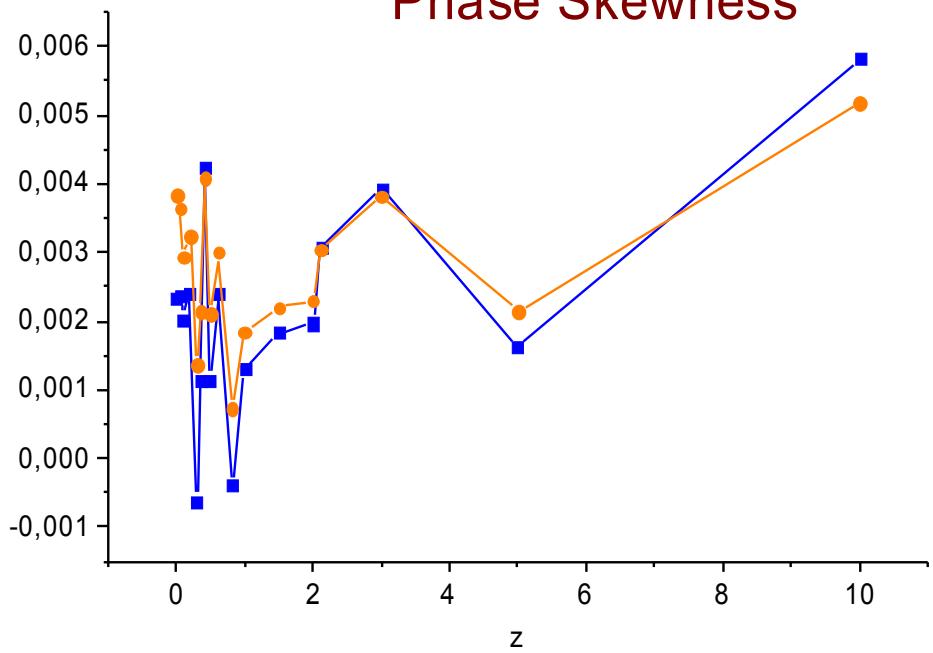




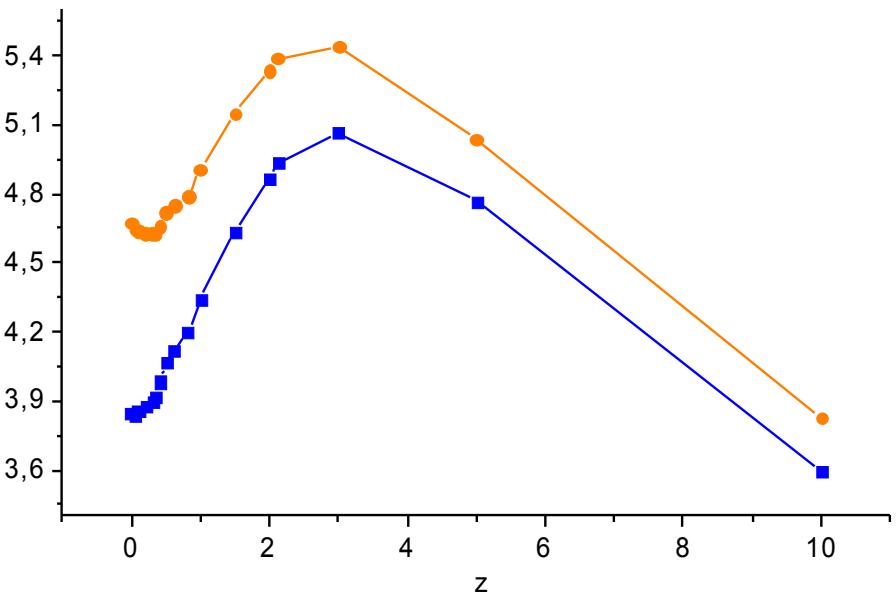
Norm Skewness



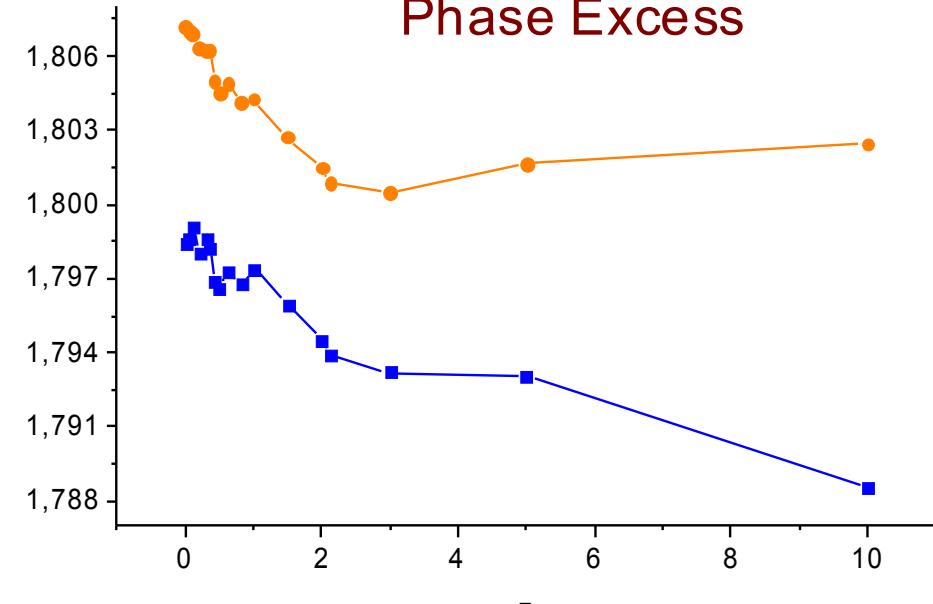
Phase Skewness

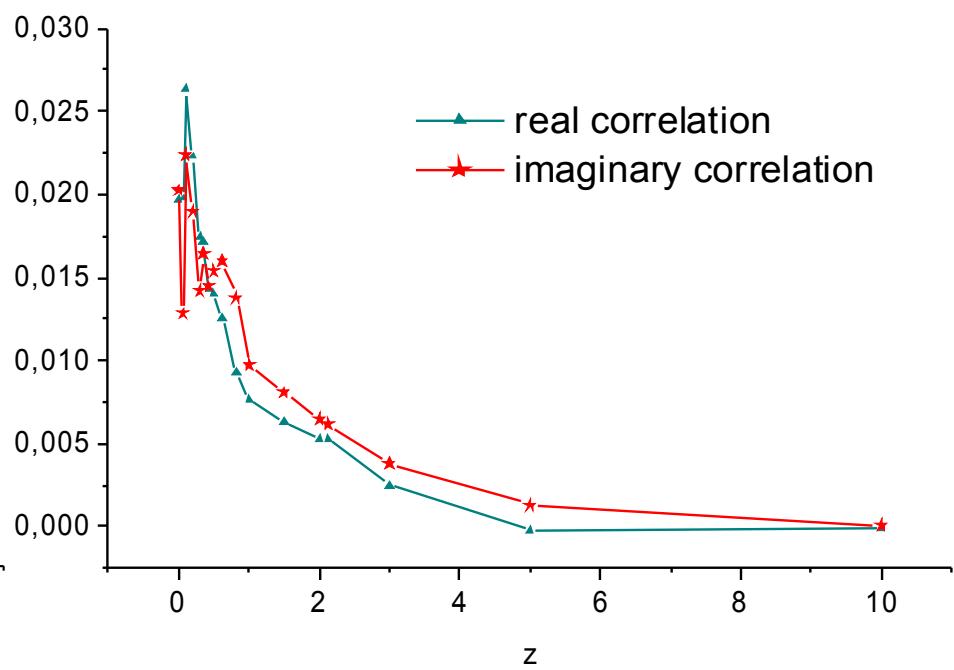
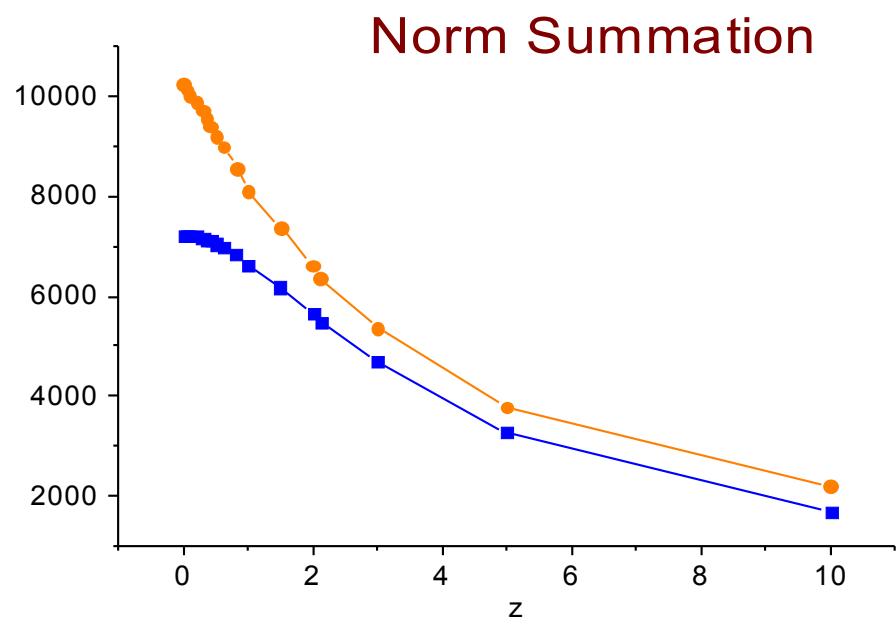
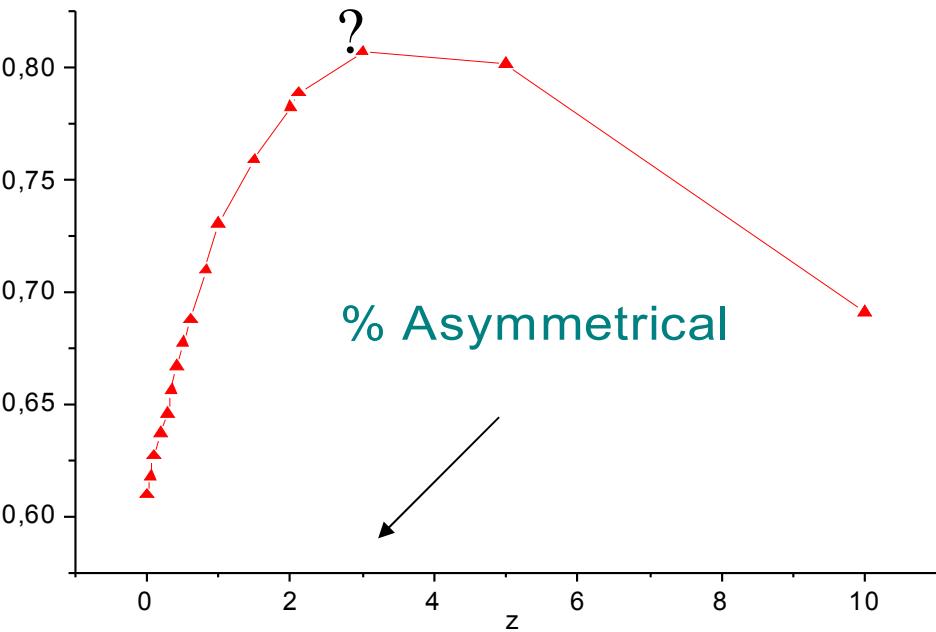
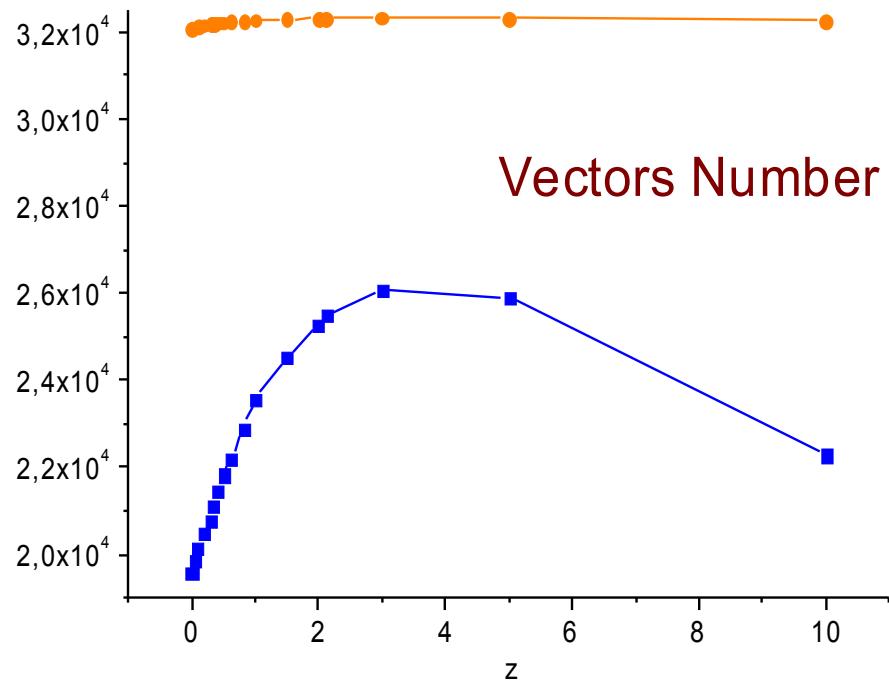


Norm Excess

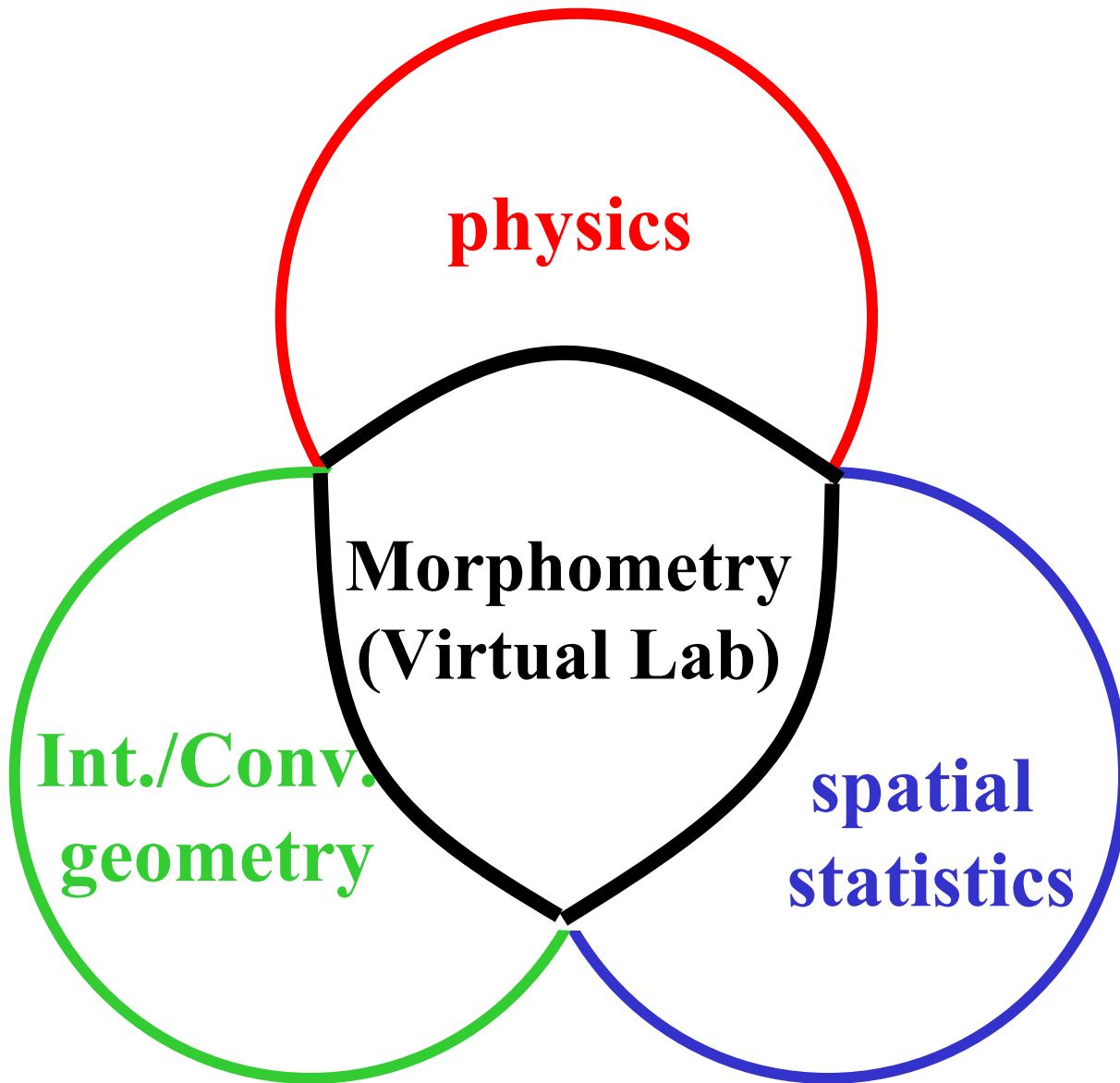


Phase Excess





Computational Morphometry for Cosmology:



PHP is a widely-used general-purpose scripting language that is especially suited for Web development and can be embedded into HTML (<http://www.php.net/>)

Pattern Analysis and Simulation projects in LAC-INPE based on *php* environment:

- 2D and 3D Analysis (GPA & MF) (ex. MHD from ZEUS)
- Dynamical 3D Analysis (a more complex usual domain)
- Solving PDEs (Burgers & KPZ)
- Analytical flyby